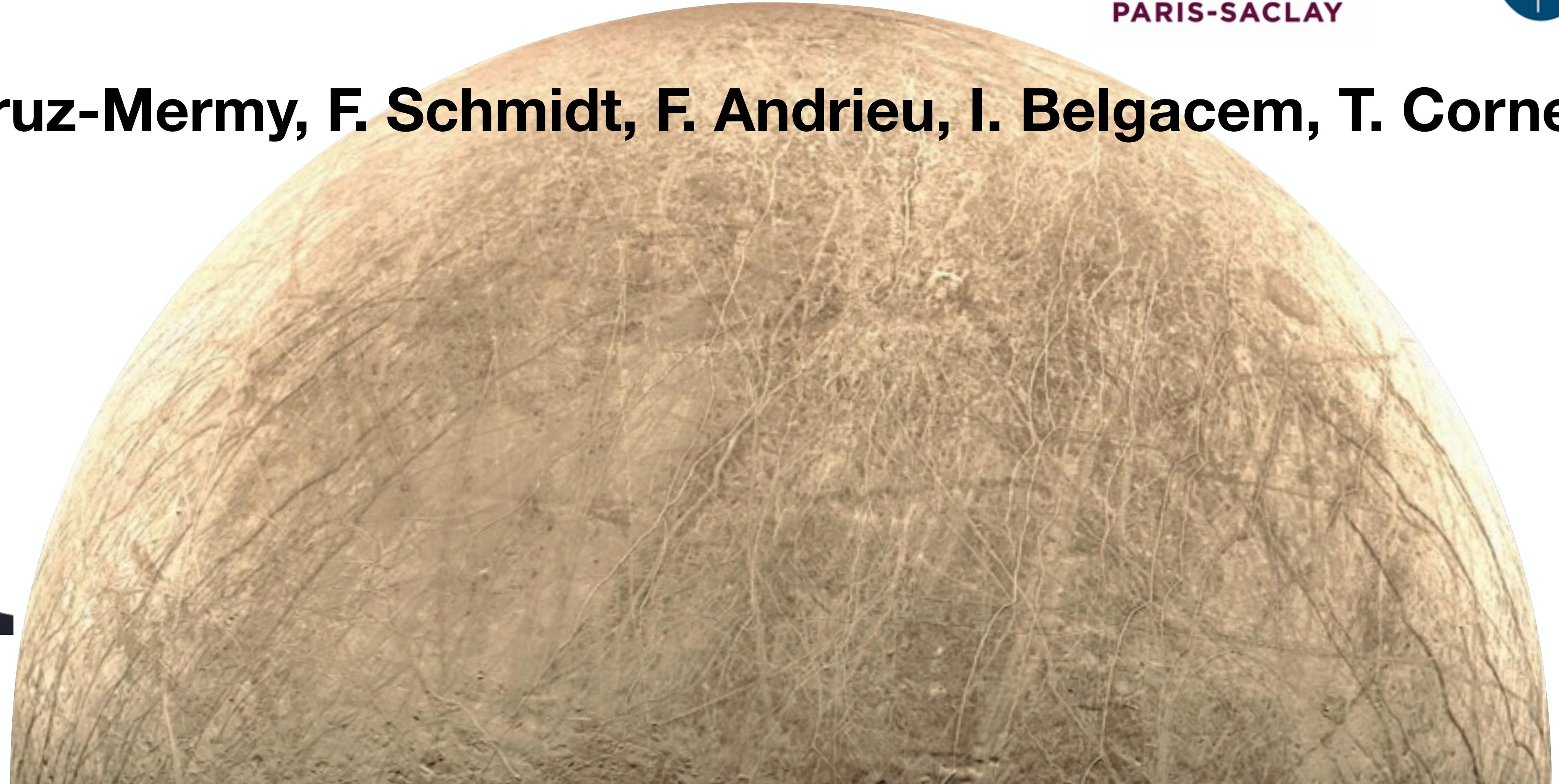


Micrometeoritics of Europa's surface using Galileo/NIMS data

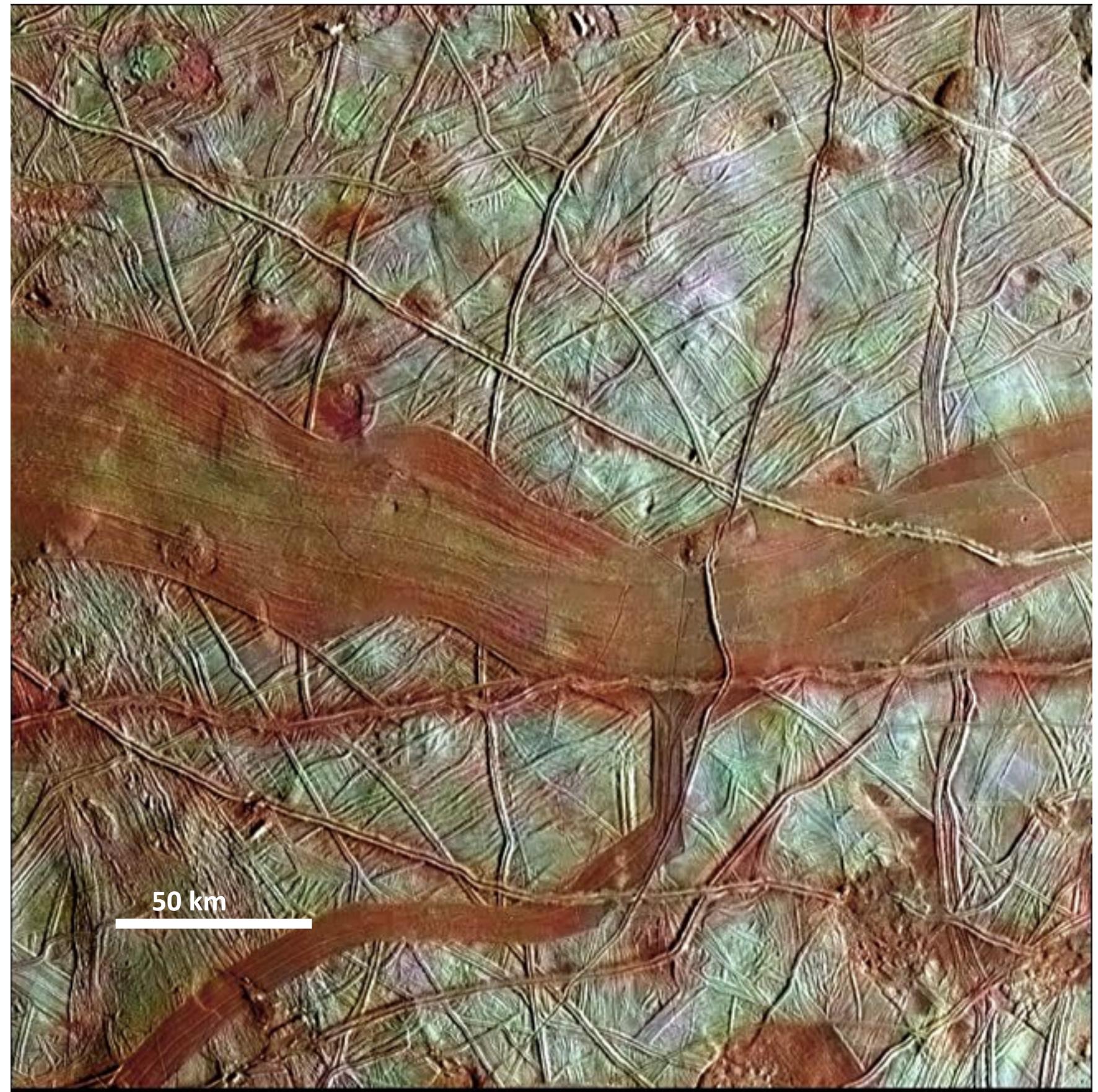


Guillaume Cruz-Mermy, F. Schmidt, F. Andrieu, I. Belgacem, T. Cornet

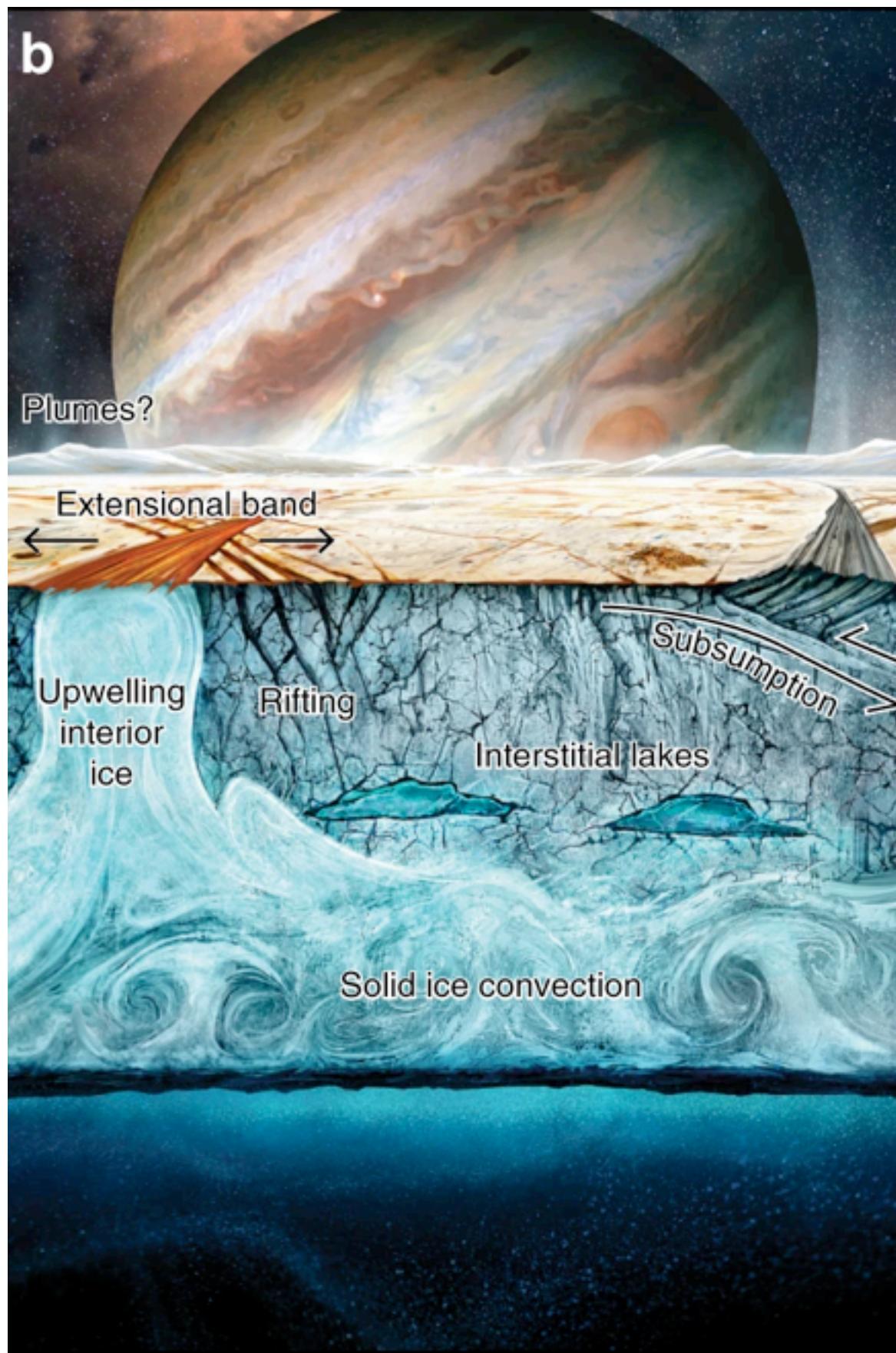


Surface morphology

Dilatation bands



Credit: NASA/JPL-Caltech/SETI Institute



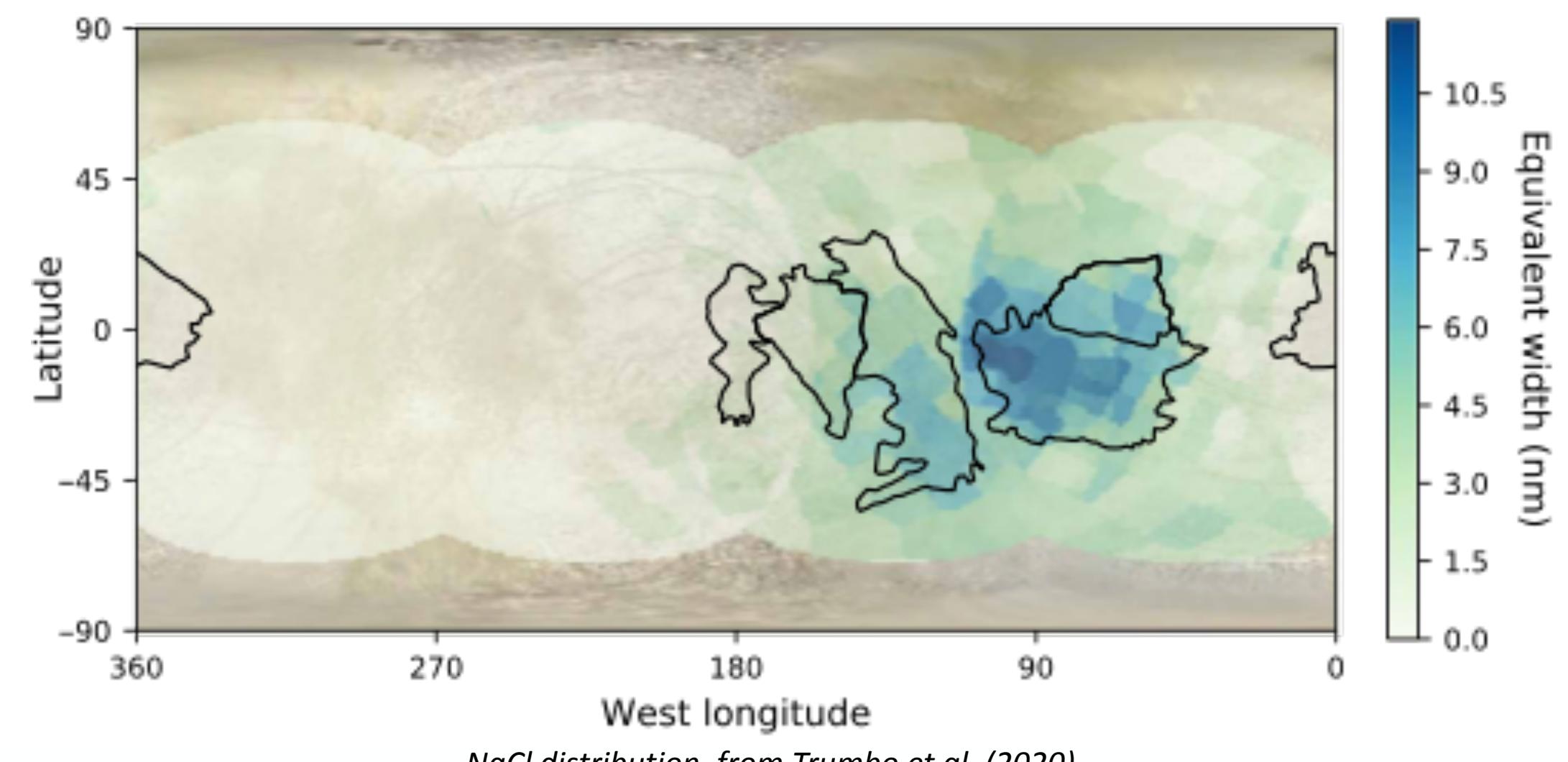
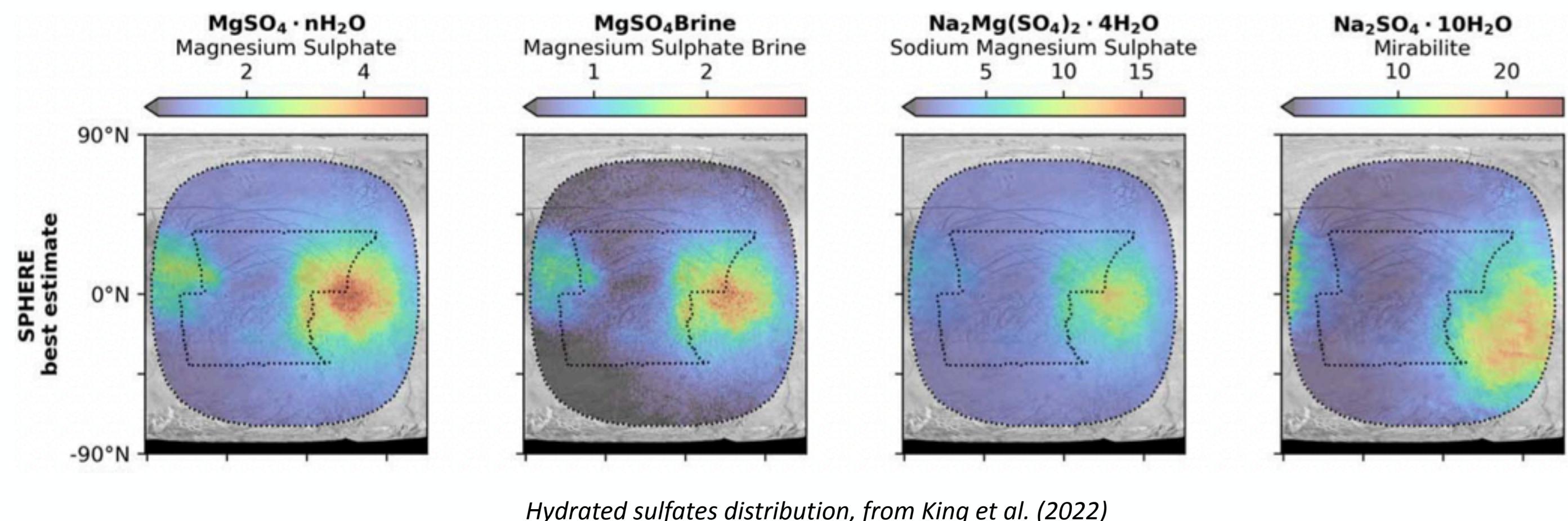
Howell & Pappalardo, 2020

- Hemispherical dichotomy (Leading Vs Trailing)
- Ridges, Linea, Bands, Craters
- Various spatial extension
- Endogenic Vs Exogenic
- Surface : key witness

Surface composition

UV-VIS-NIR Spectroscopy

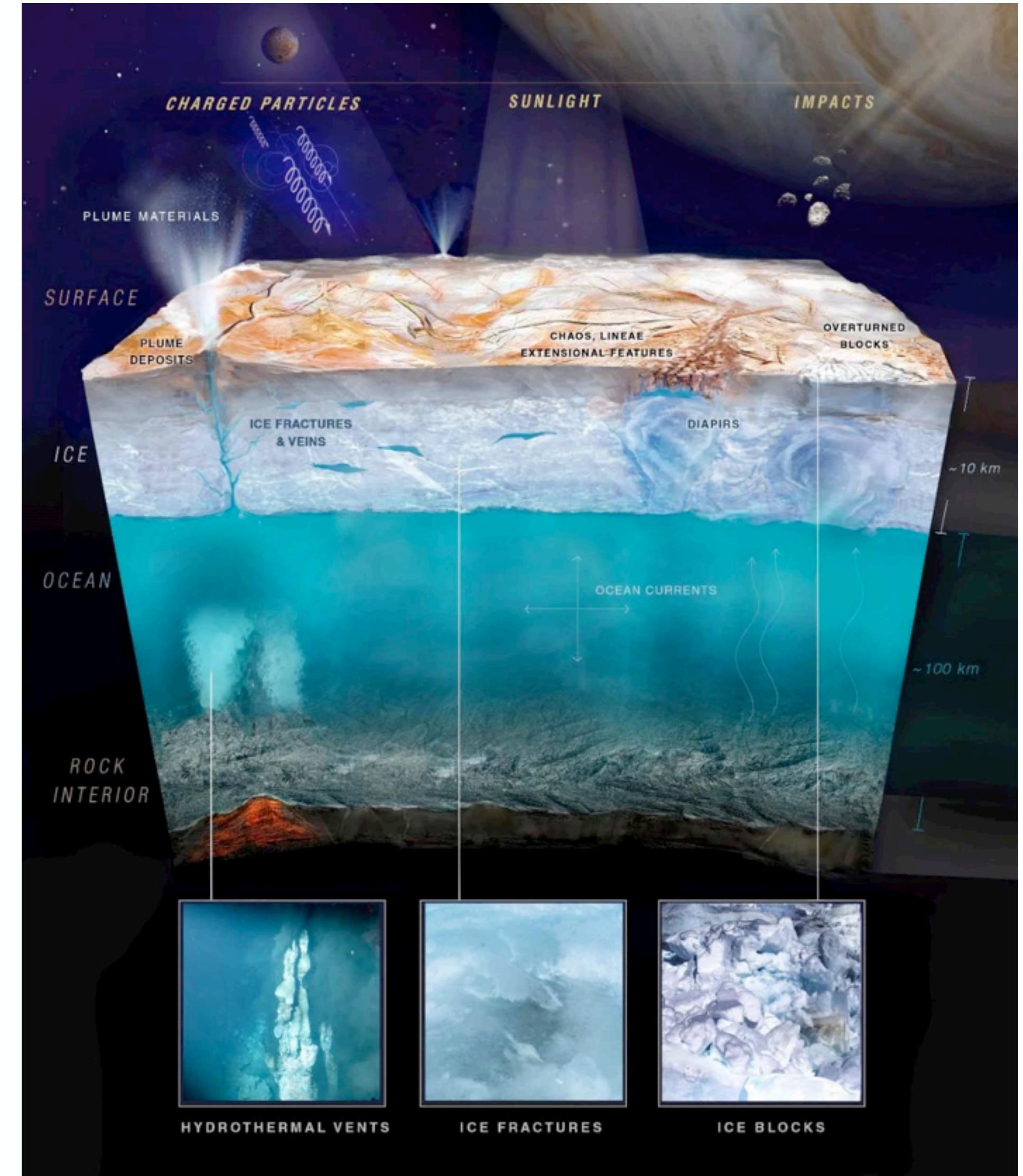
- Water ice (crystalline/amorphous)
 - Kuiper et al. (1957), Calvin et al. (1995)
- Sulfuric Acid Octahydrate (SAO)
 - Carlson et al. (1999, 2005), Mishra et al. (2021)
- Hydrated sulfates
 - McCord et al. (1998), Dalton et al. (2007, 2012), Ligier et al. (2016), King et al. (2022)
- Chlorinates
 - Brown et al. (2013), Hanley et al. (2014), Ligier et al. (2016), Trumbo et al. (2020)
- Oxidants
 - Carlson et al. (1999), Hand et al. (2006), Trumbo et al. (2019)



Why studying the surface ?

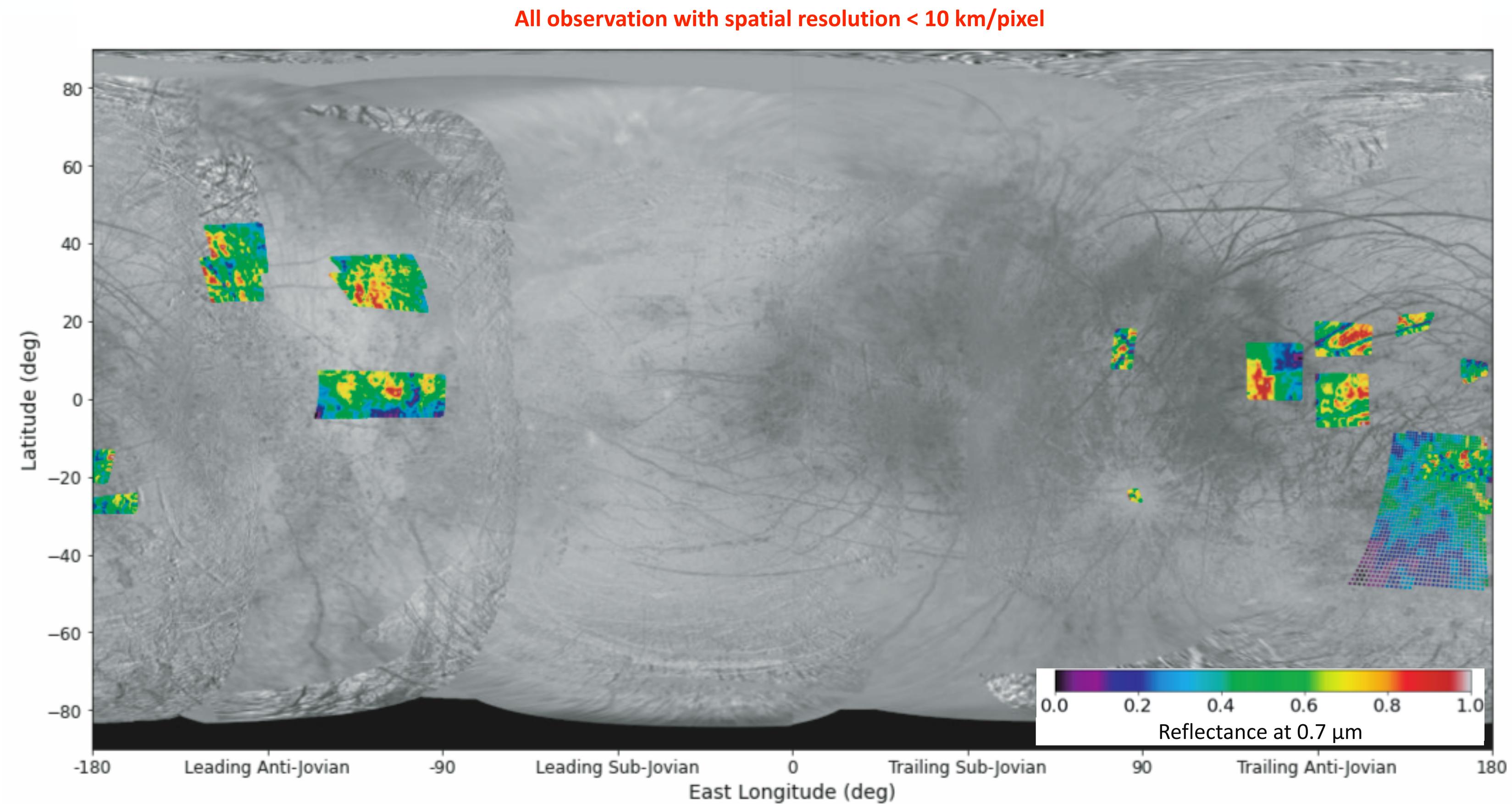
- **Previous studies:**
 - Linear mixture model / Hapke model
 - Spectral comparison
 - Low spatial resolution observations
- **Objectives:**
 - Use accurate radiative transfer & bayesian approach
 - Use Galileo/NIMS high spatial resolution images
 - Differentiate between endogenous and exogenous processes

**What is the true surface composition ?
What is the ice microphysics ?**



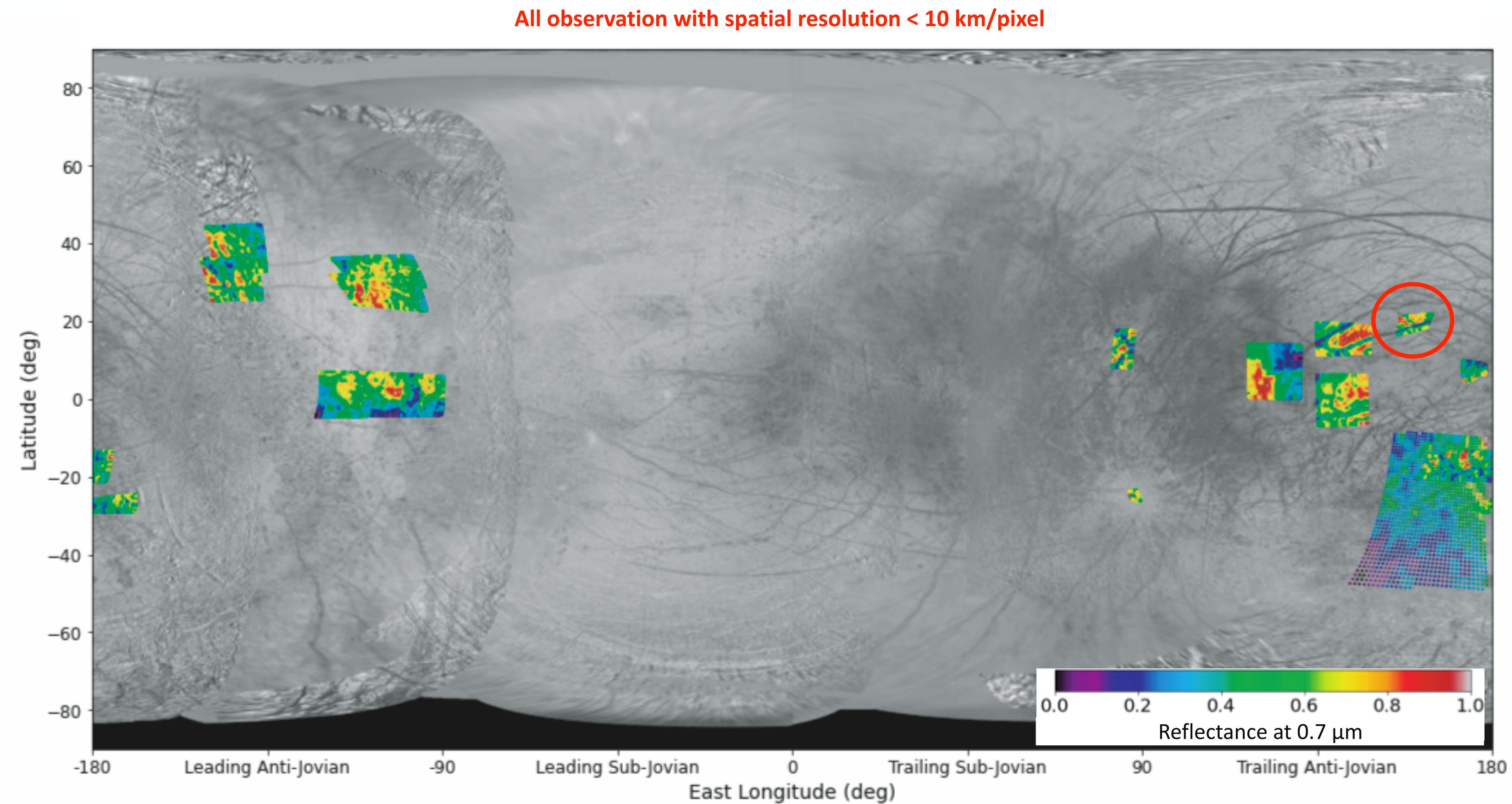
Galileo/NIMS hyperspectral data

- NIR wavelengths:
0.7 - 5.2 μm
- Spectral resolution:
0.125 - 0.250 μm
- Spatial resolution up
to **2.5 km/pixel** (flyby)

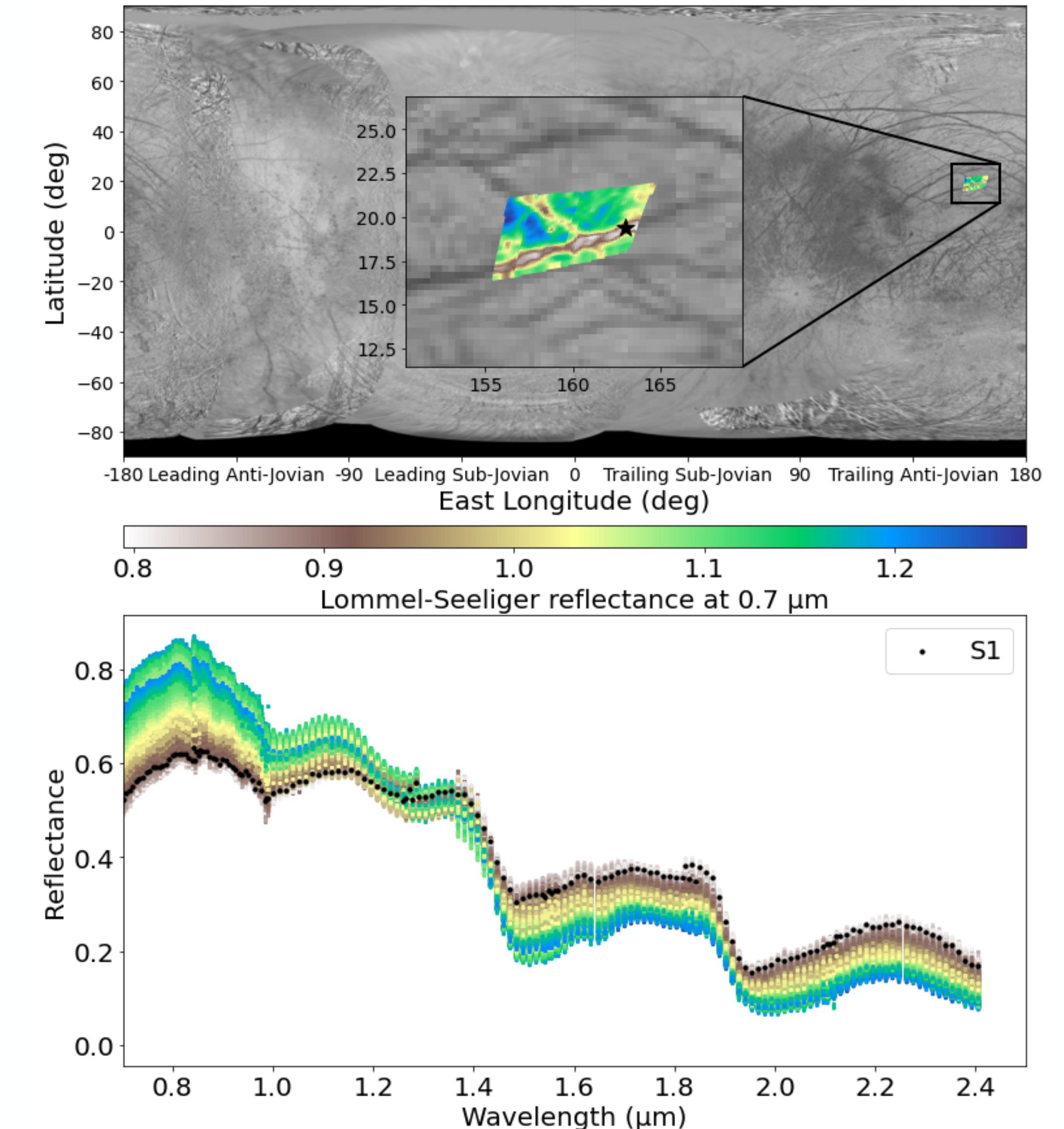


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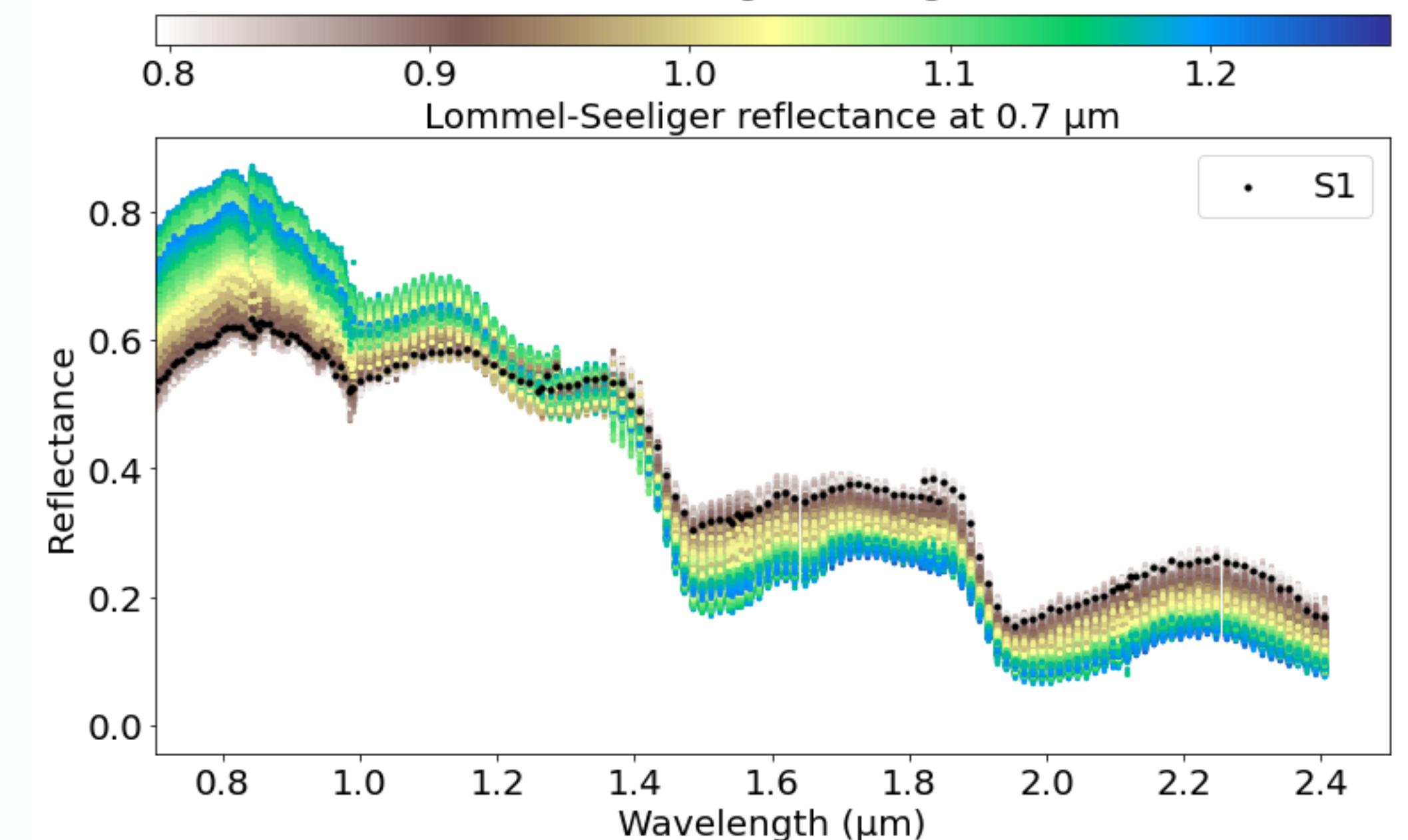
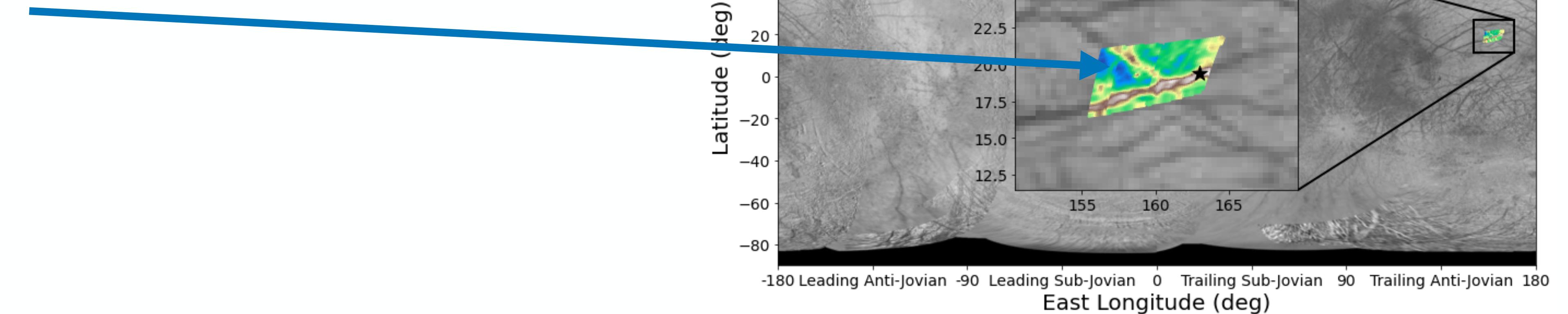


Test case: Harmonia Linea



Test case: Harmonia Linea

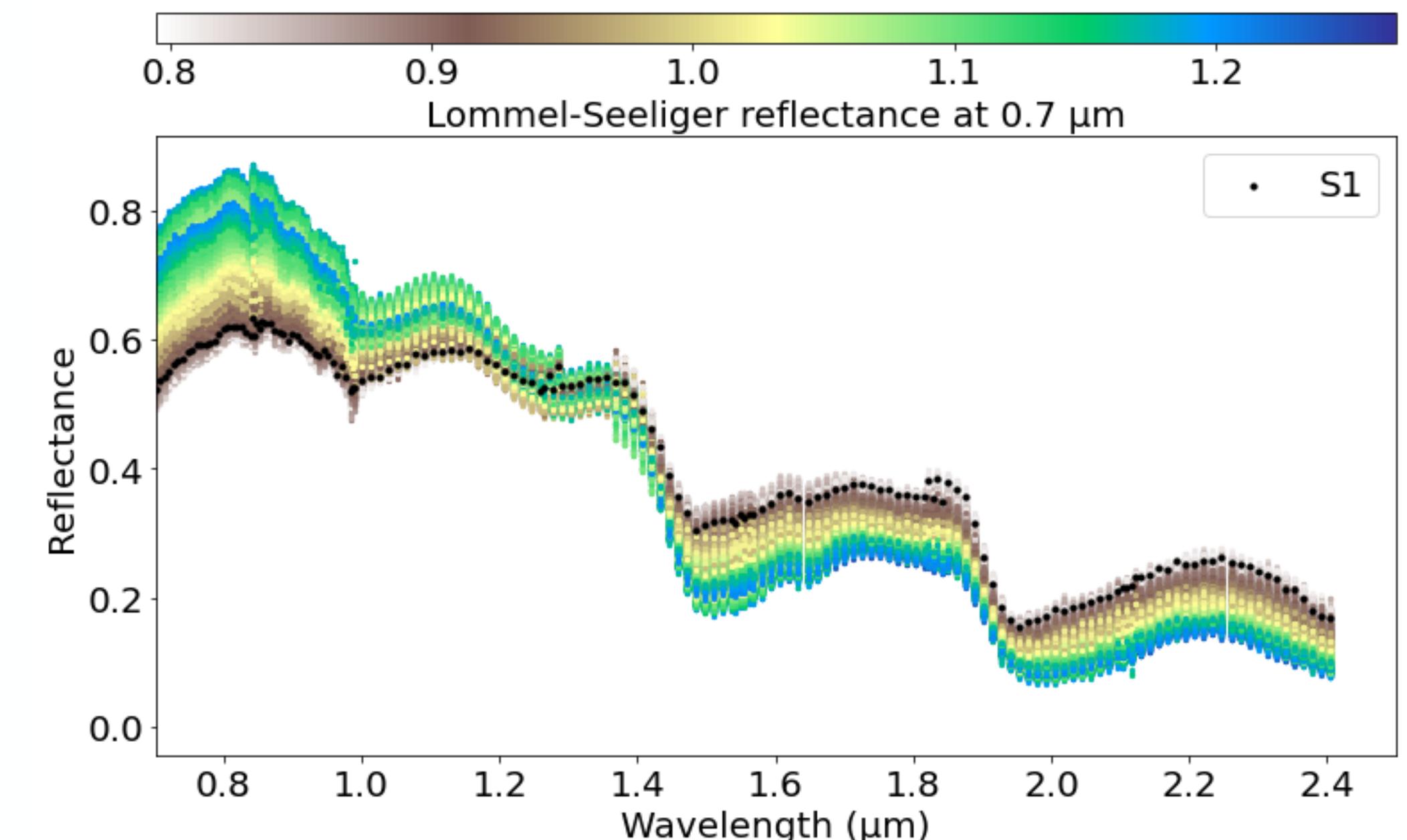
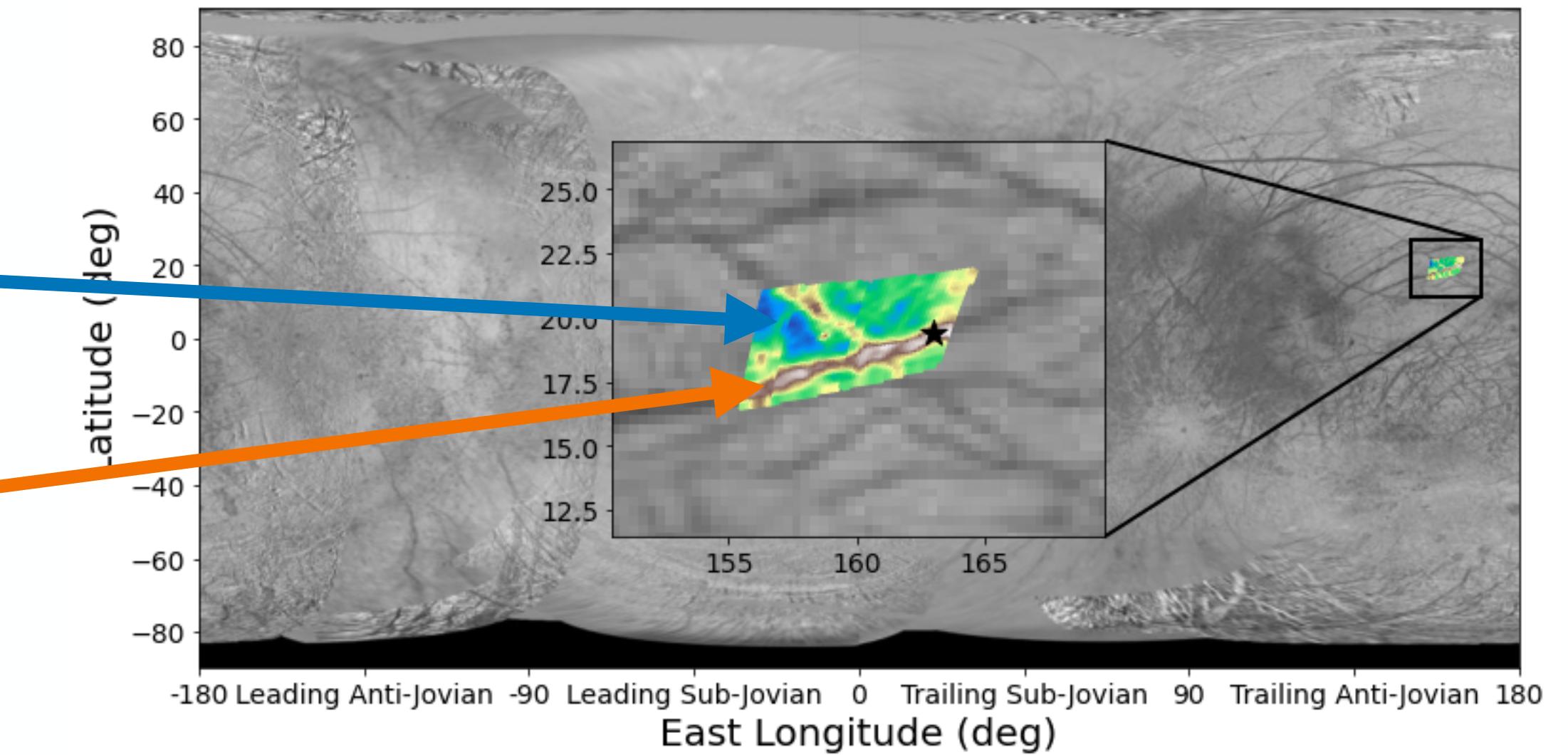
Bright plains



Test case: Harmonia Linea

Bright plains

Dark lineaments



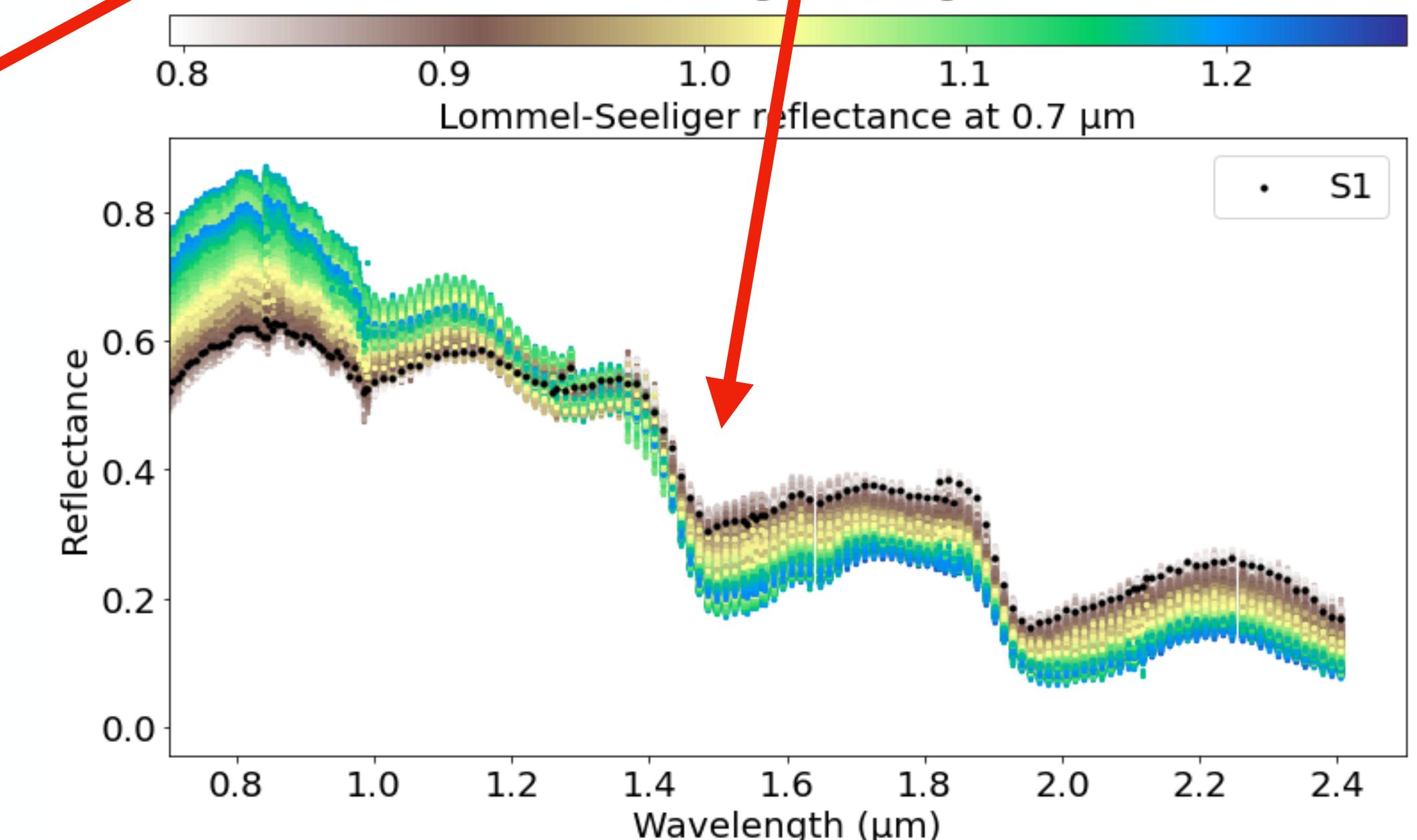
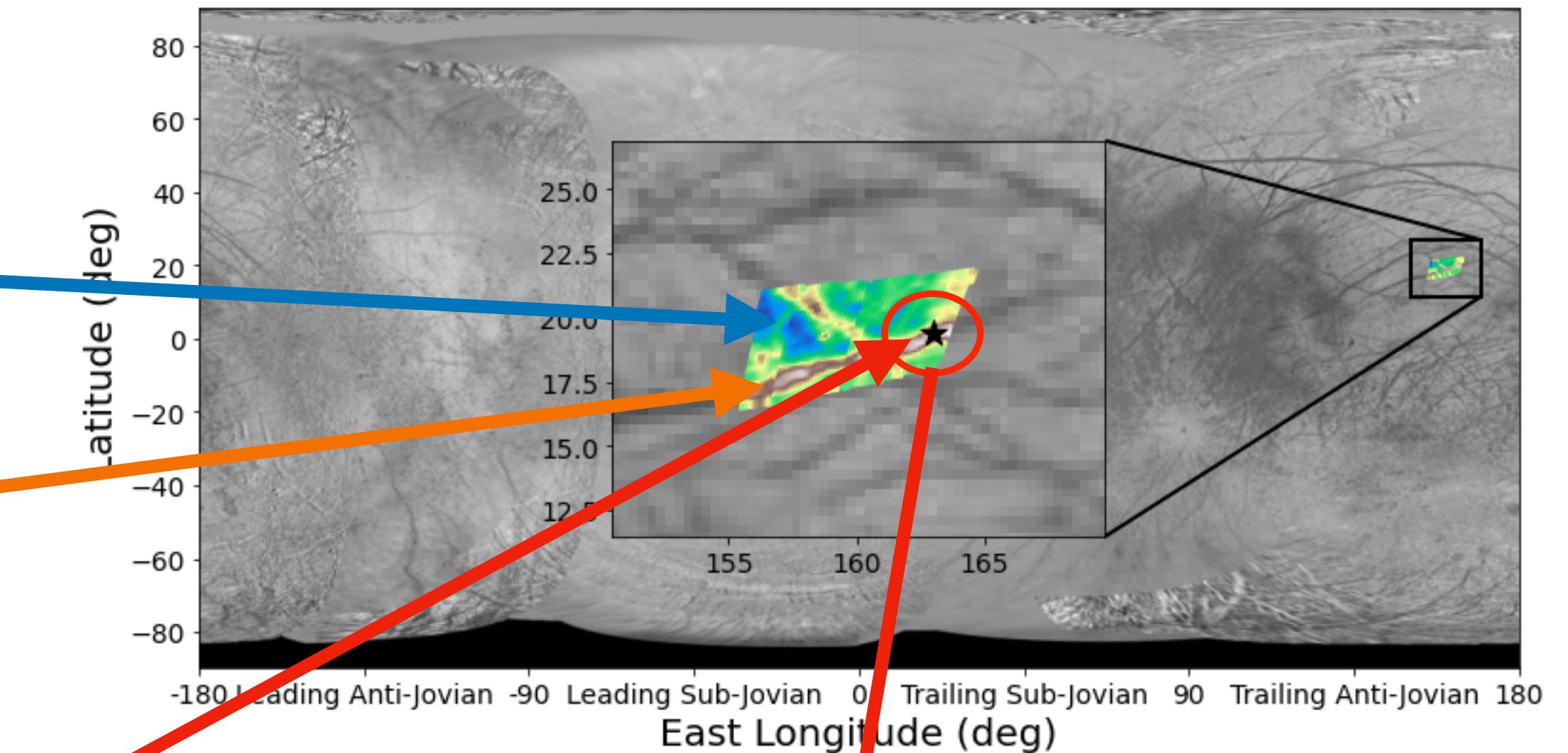
Test case: Harmonia Linea

Bright plains

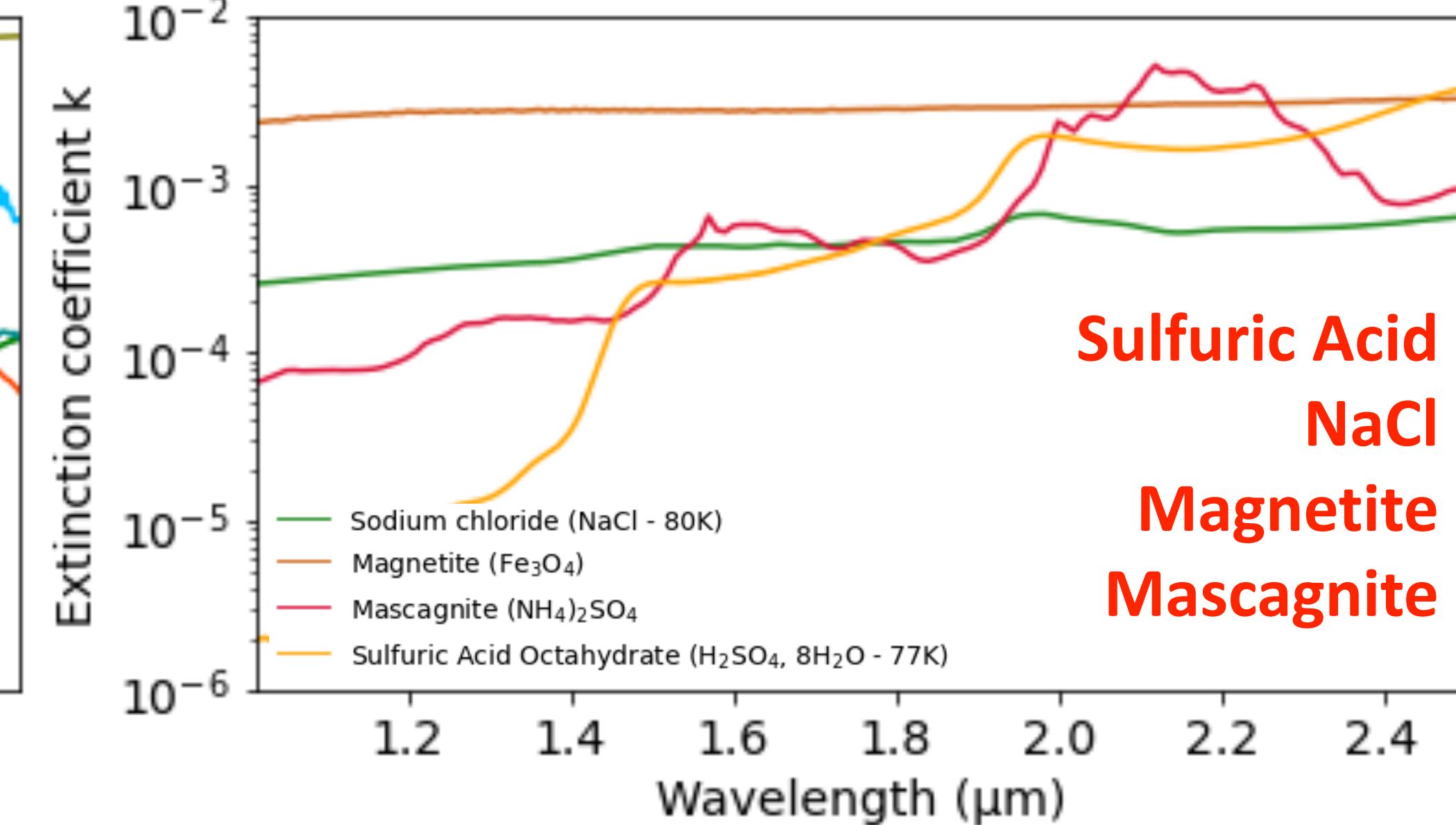
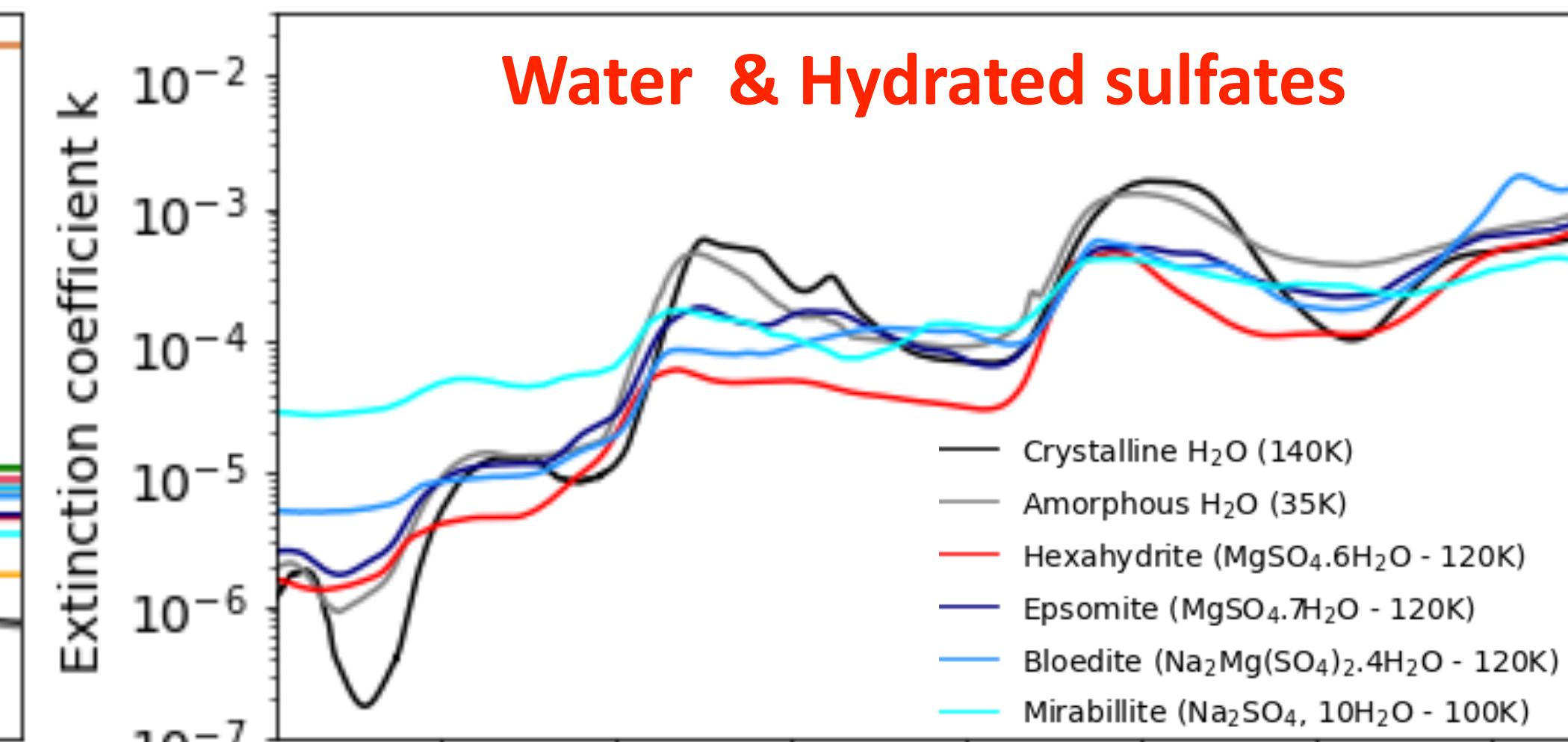
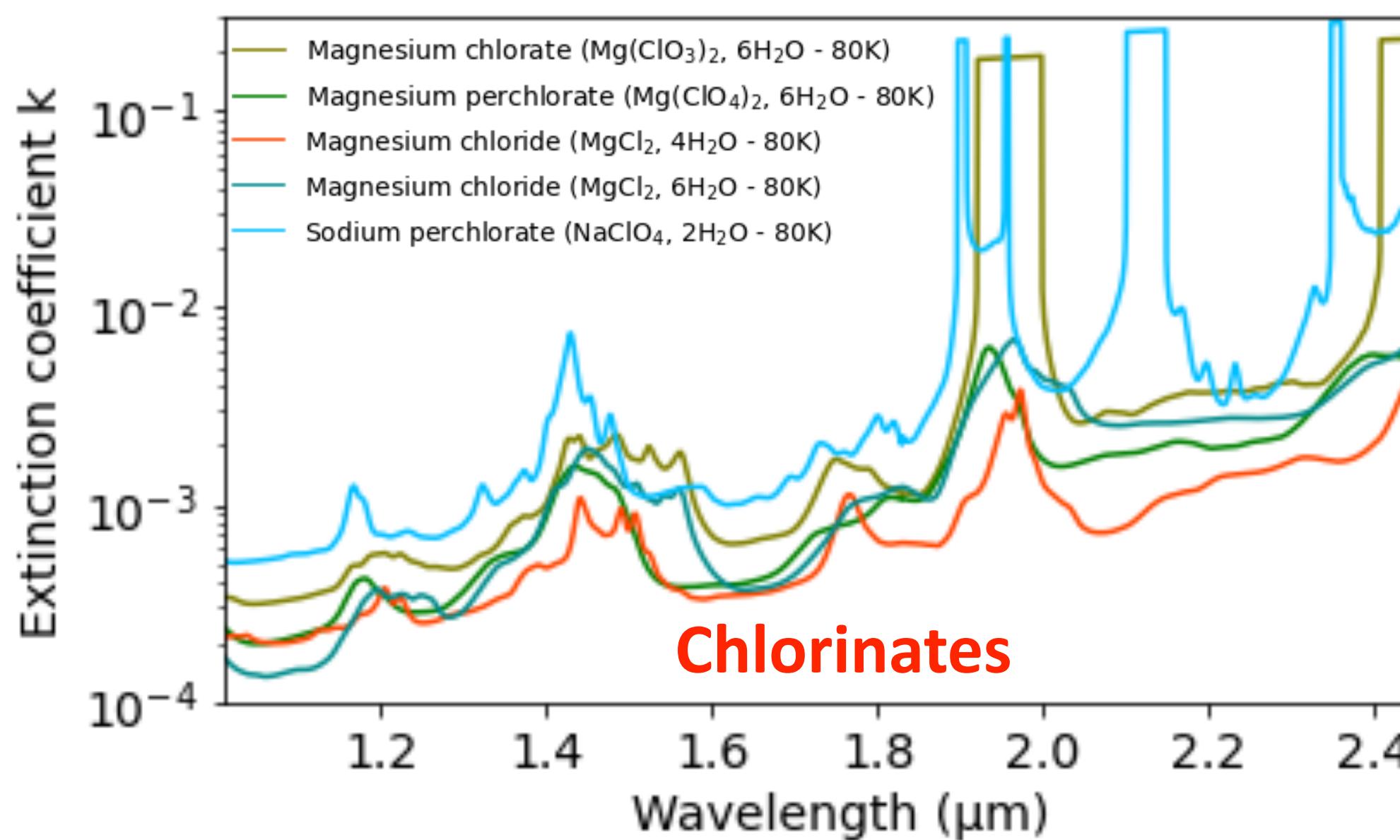
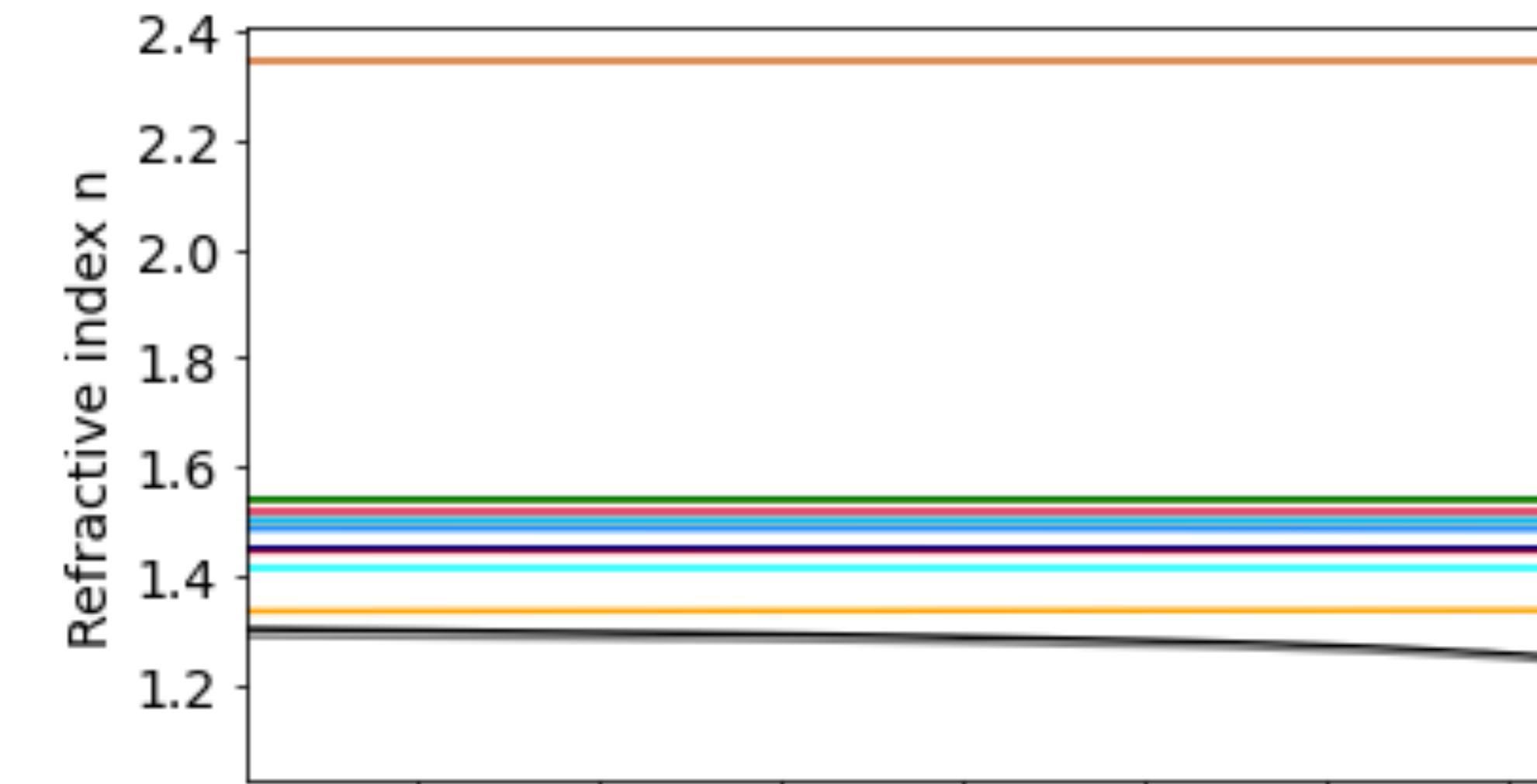
Dark lineaments

Reference spectrum

- Good spectral res.
- Distorted absorption bands
- Hard to fit !



Compound selection

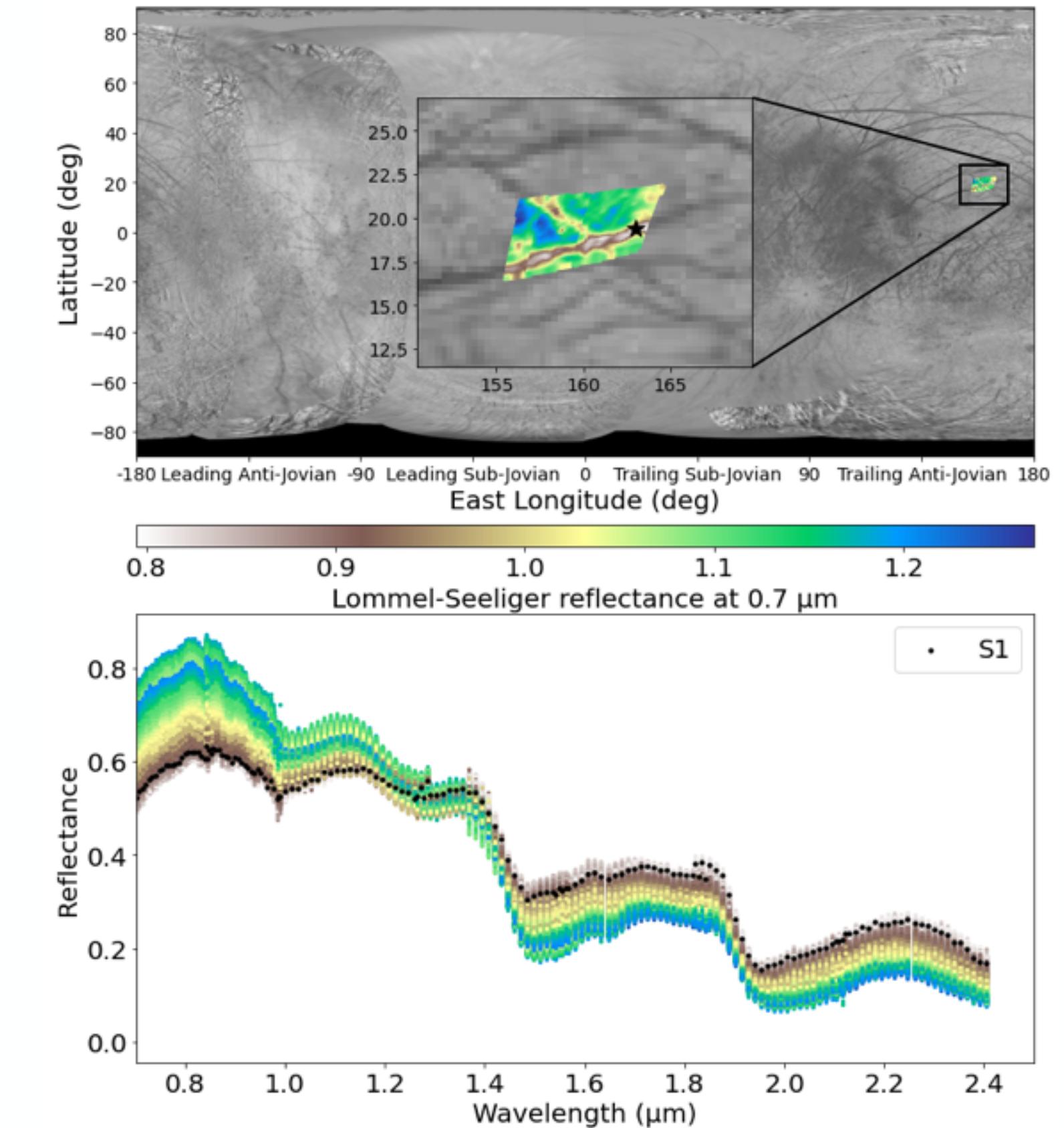


Compound selection

- K out of 15 potential endmembers

— Crystalline H ₂ O (140K)	— Magnesium chloride (MgCl ₂ , 4H ₂ O - 80K)
— Amorphous H ₂ O (35K)	— Magnesium chloride (MgCl ₂ , 6H ₂ O - 80K)
— Hexahydrite (MgSO ₄ .6H ₂ O - 120K)	— Sodium perchlorate (NaClO ₄ , 2H ₂ O - 80K)
— Epsomite (MgSO ₄ .7H ₂ O - 120K)	— Sodium chloride (NaCl - 80K)
— Bloedite (Na ₂ Mg(SO ₄) ₂ .4H ₂ O - 120K)	— Magnetite (Fe ₃ O ₄)
— Mirabilite (Na ₂ SO ₄ , 10H ₂ O - 100K)	— Mascagnite (NH ₄) ₂ SO ₄
— Magnesium chlorate (Mg(ClO ₃) ₂ , 6H ₂ O - 80K)	— Sulfuric Acid Octahydrate (H ₂ SO ₄ , 8H ₂ O - 77K)
— Magnesium perchlorate (Mg(ClO ₄) ₂ , 6H ₂ O - 80K)	

- Radiative transfer: Hapke et al. (1993, 2012)
 - Volume abundance, Grain size, Surface roughness
- Bayesian MCMC
 - DEMCz algorithm (Python mc3): Cubillos et al. (2016)



$$\binom{n}{k} = \frac{n!}{k! \times (n-k)!}$$

455,
1365 and 3003
combinations

k = 3, 4 and 5

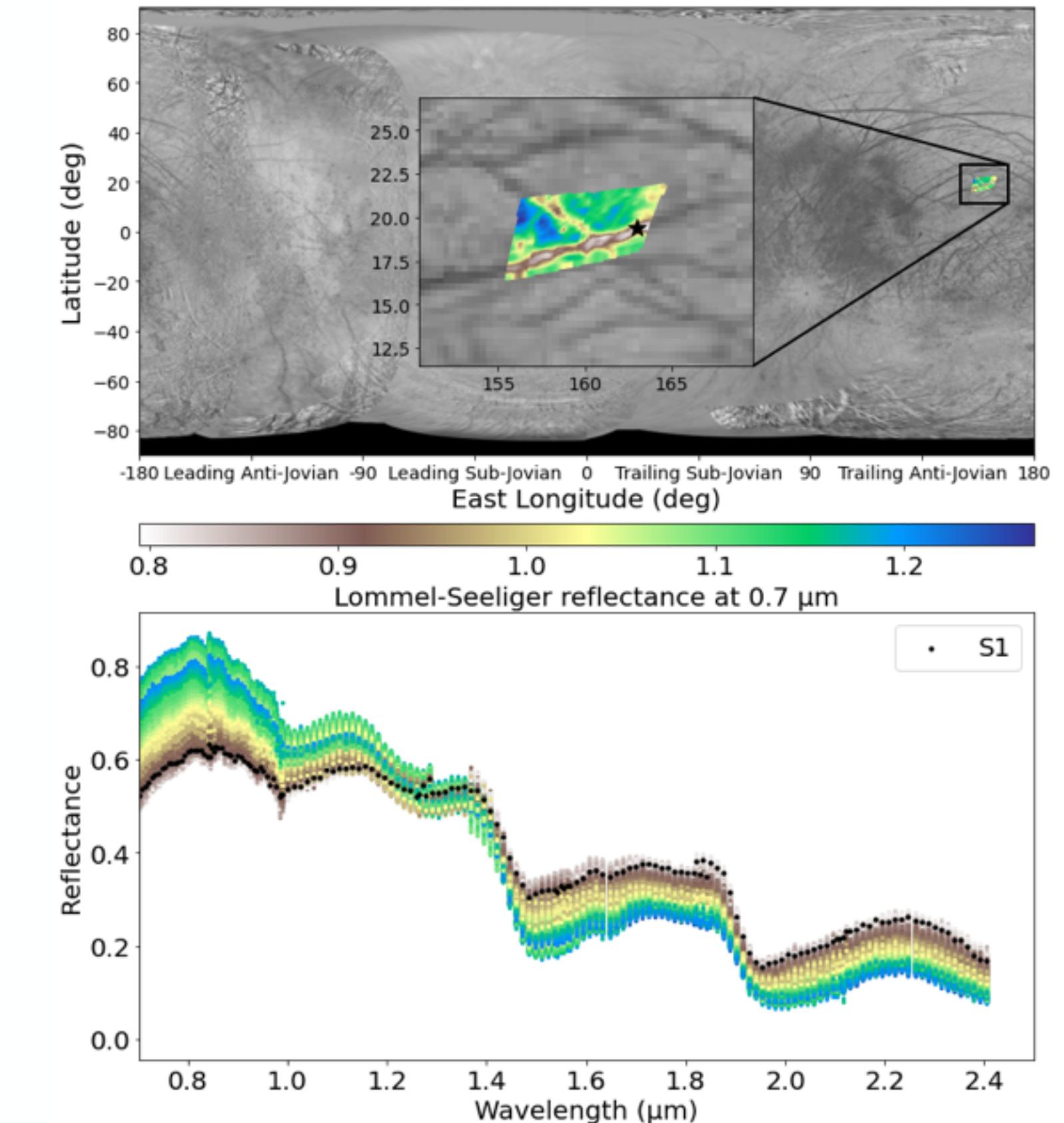
Which compounds to use ? Test all combinations !

Compound selection

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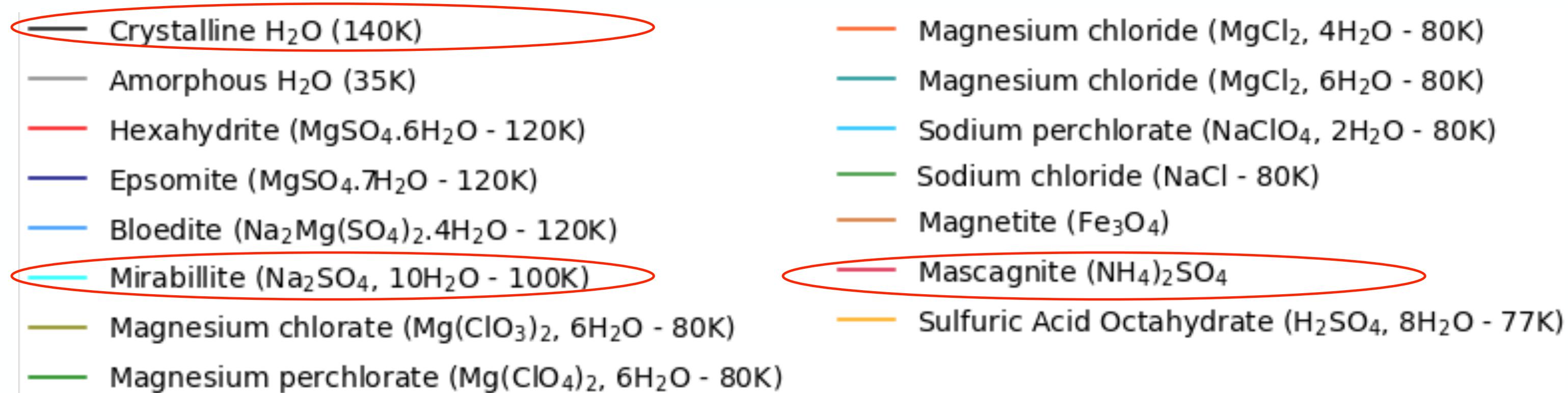
→ 455,
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$k = 3, 4 \text{ and } 5$

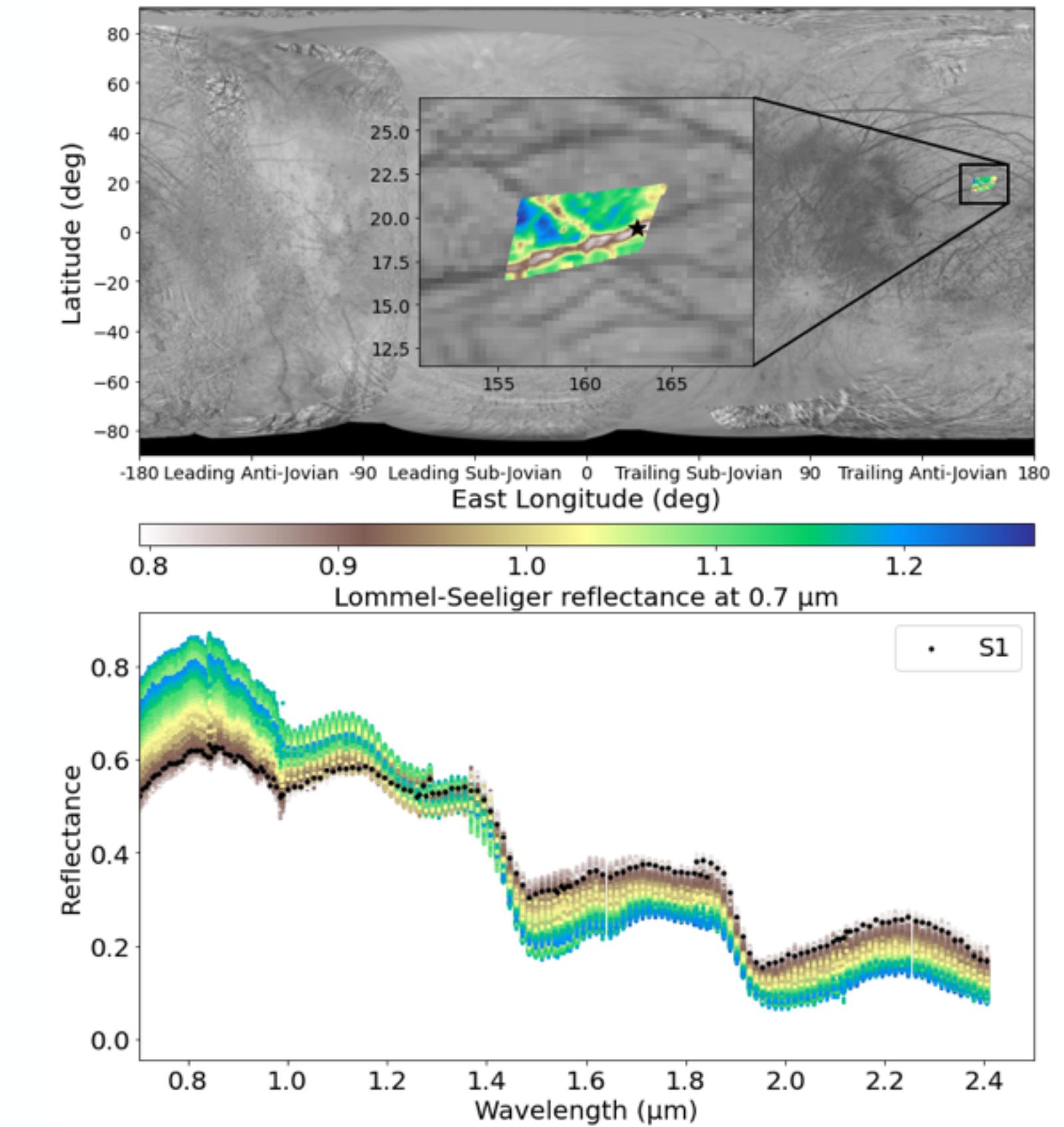
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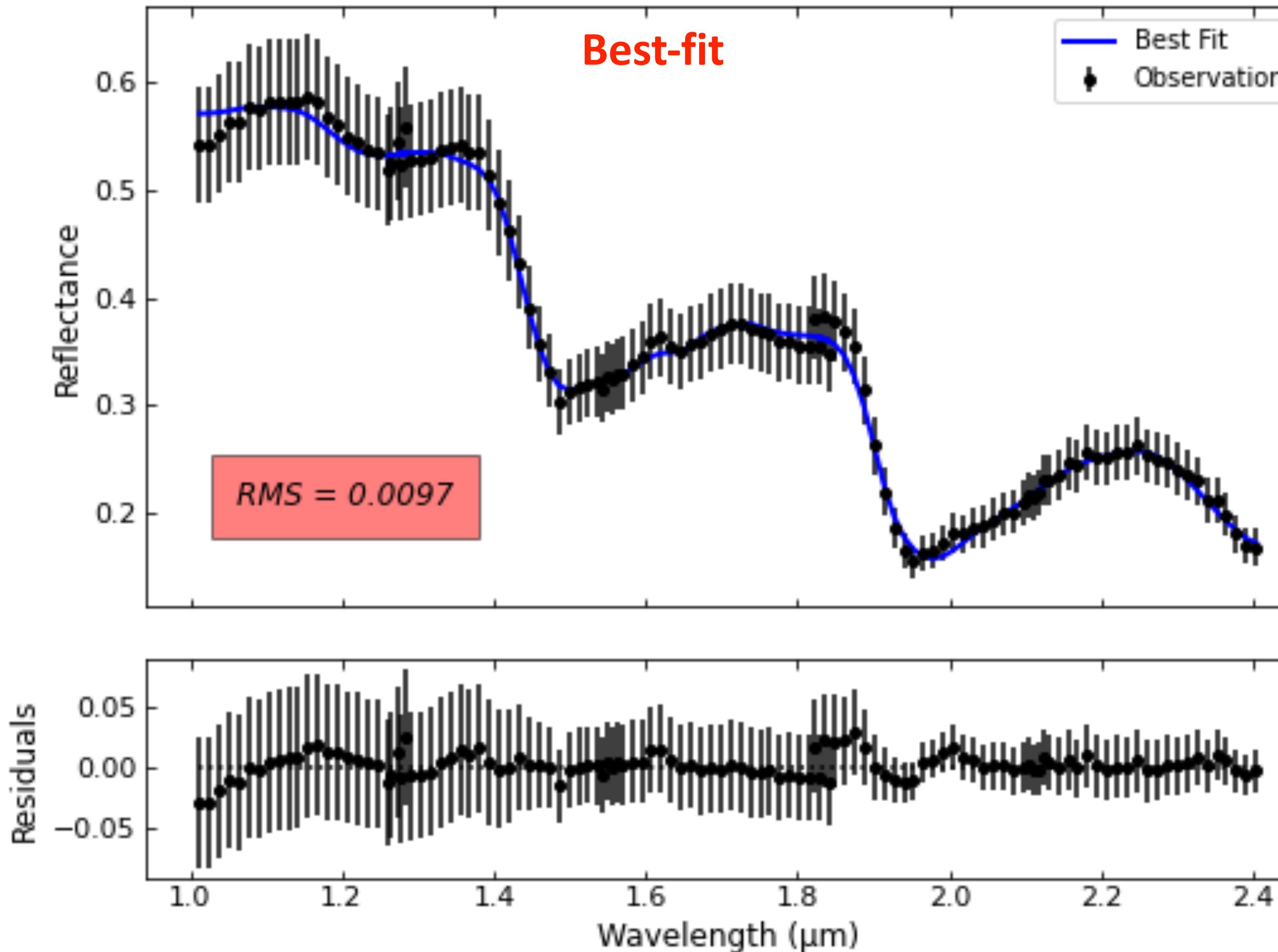
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→ 455,
1365 and 3003
combinations

$k = 3, 4 \text{ and } 5$

Which compounds to use ? Test all combinations !

Typical fit

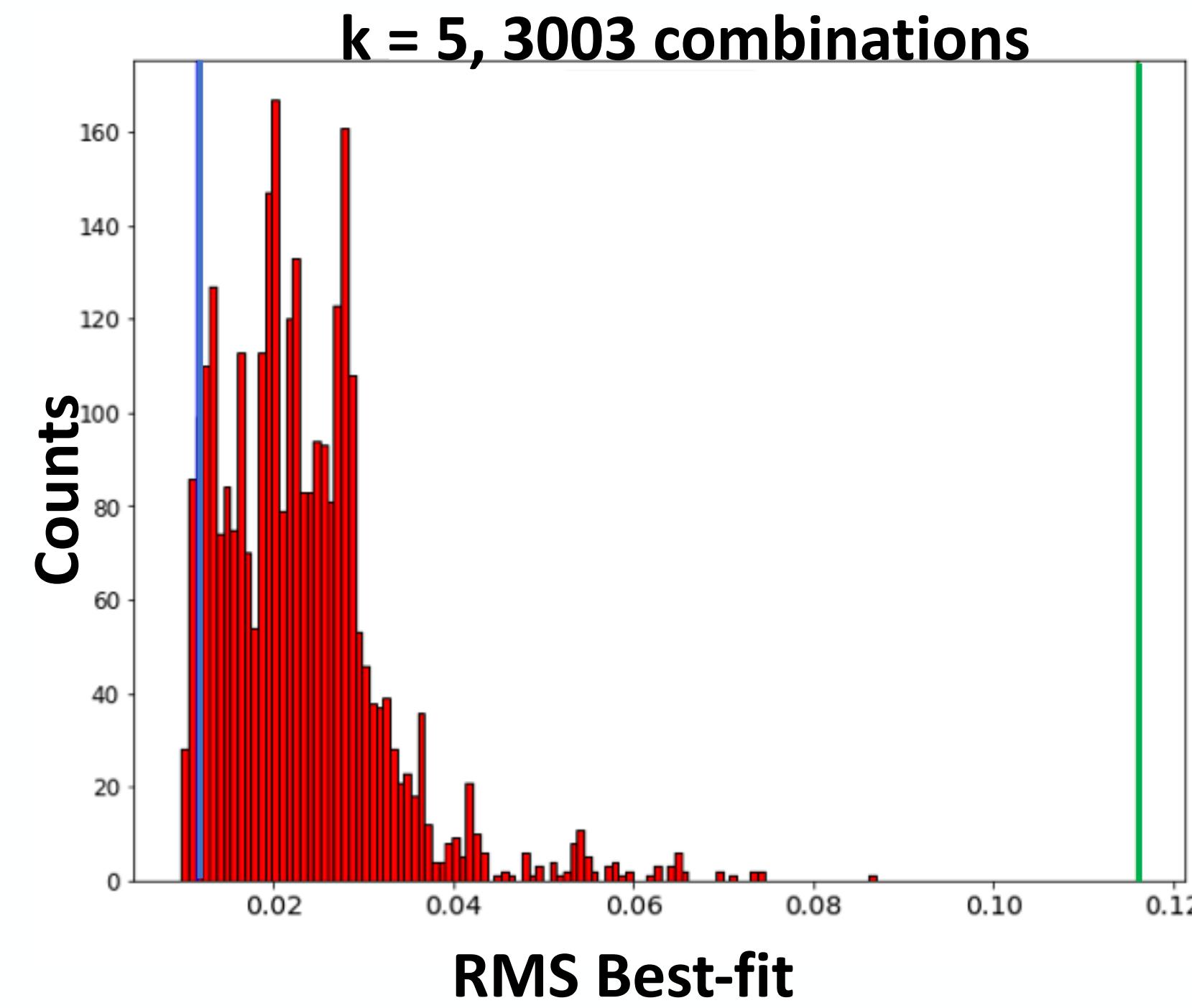
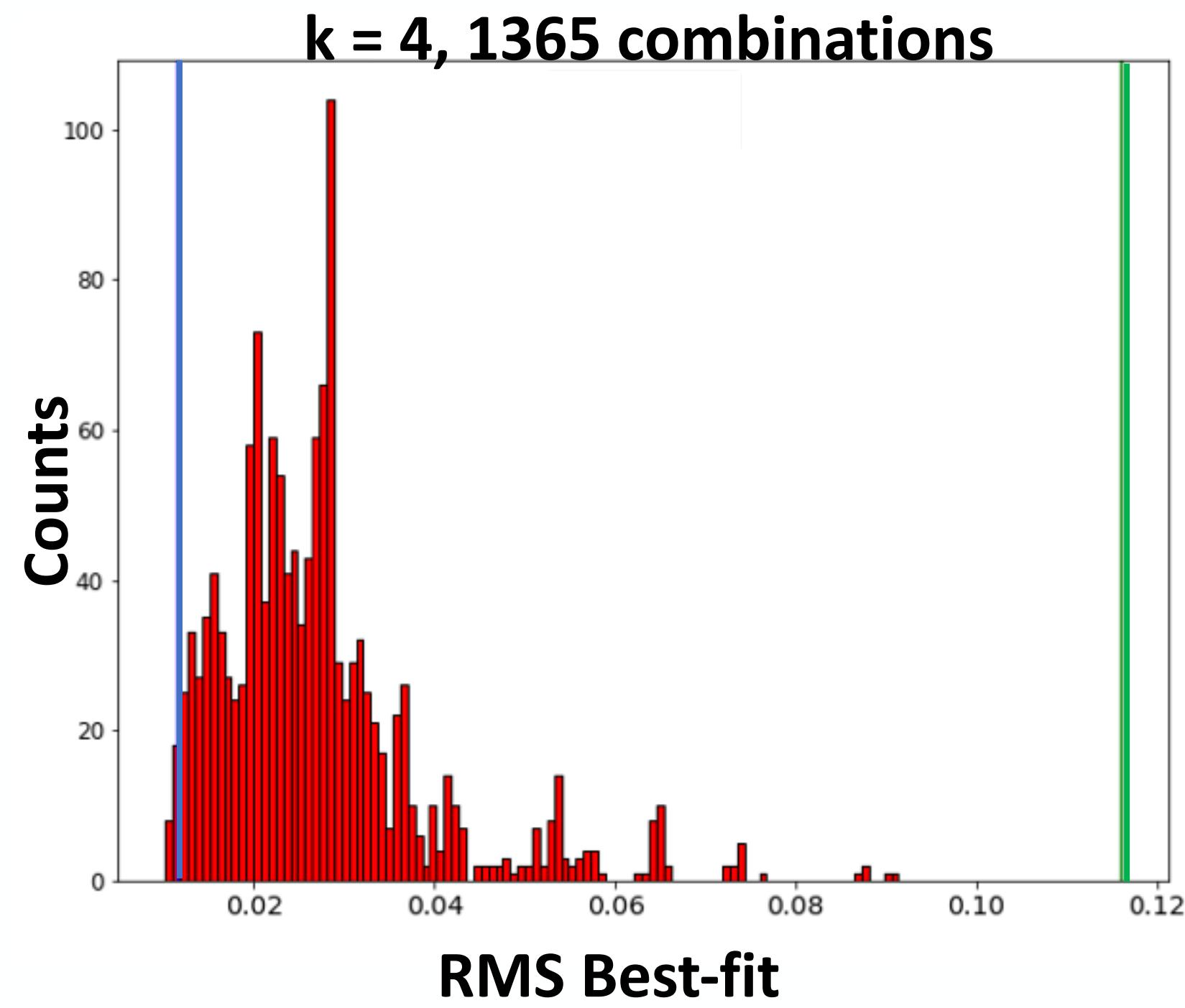
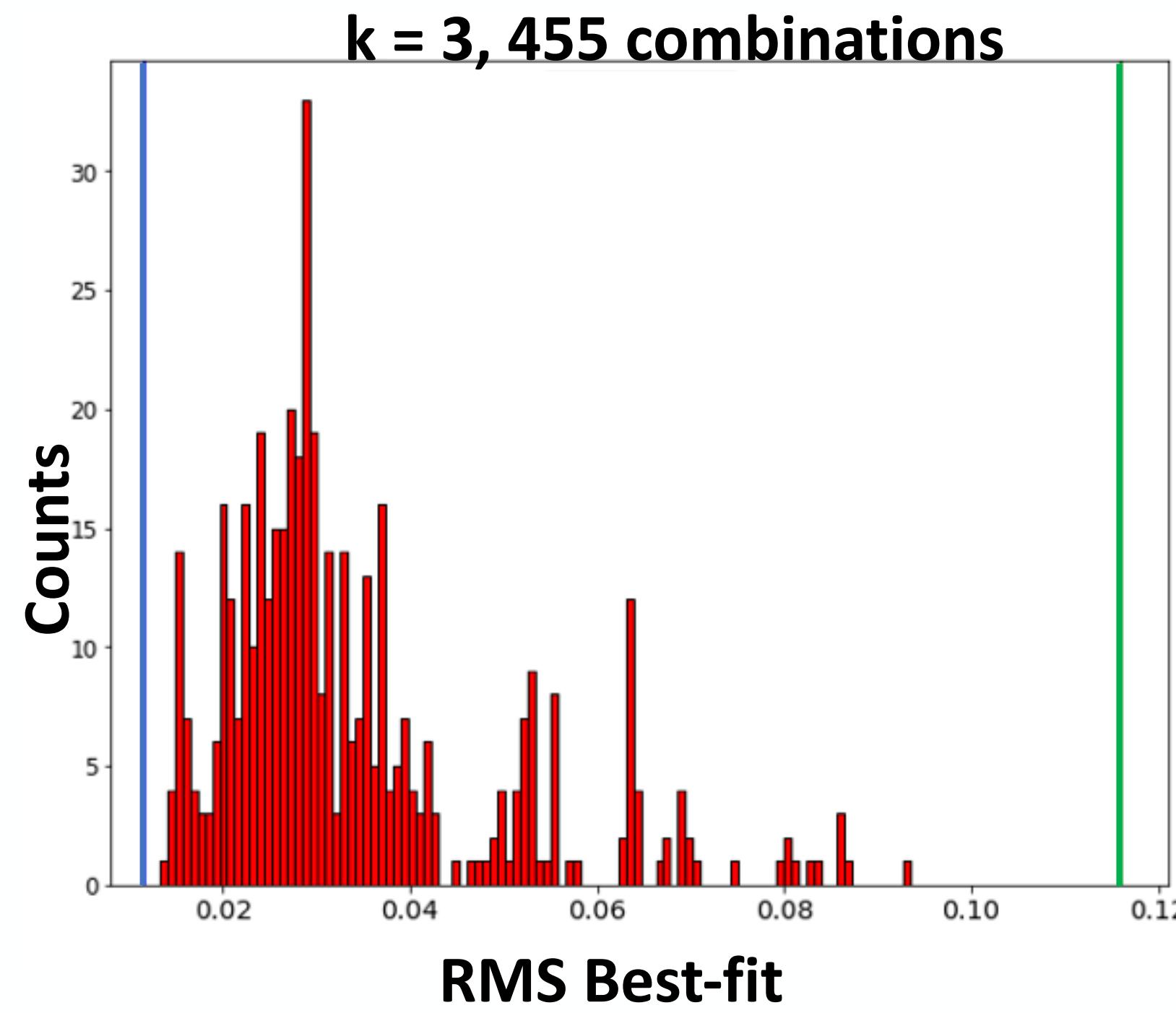


H_2O
 $\text{MgSO}_4 \cdot 6\text{H}_2\text{O}$
 $\text{H}_2\text{SO}_4 \cdot 8\text{H}_2\text{O}$
 $\text{MgCl}_2 \cdot 4\text{H}_2\text{O}$
 $\text{NaSO}_4 \cdot 10\text{H}_2\text{O}$

- Data uncertainty:
10% (Carlson et al., 1992)

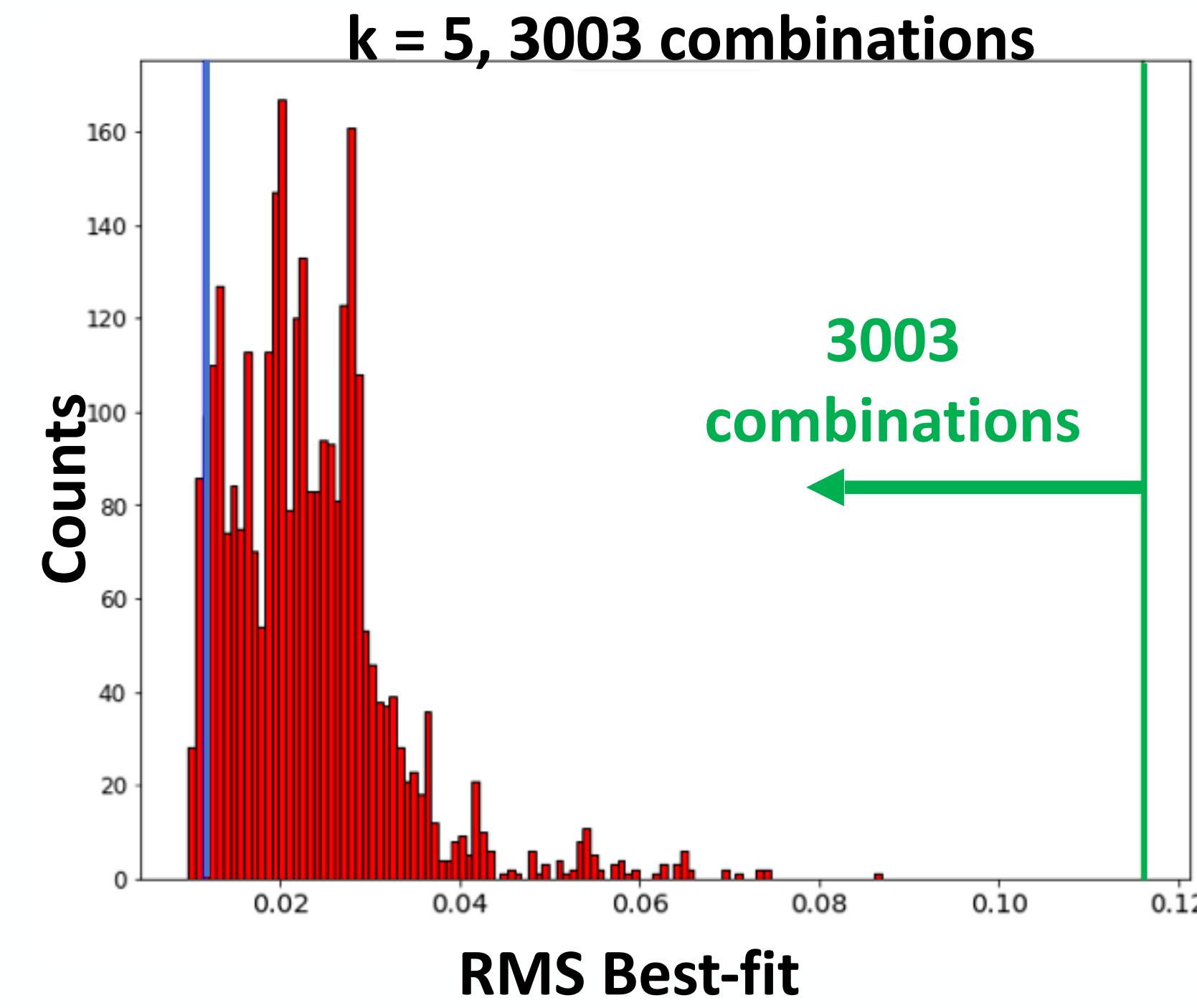
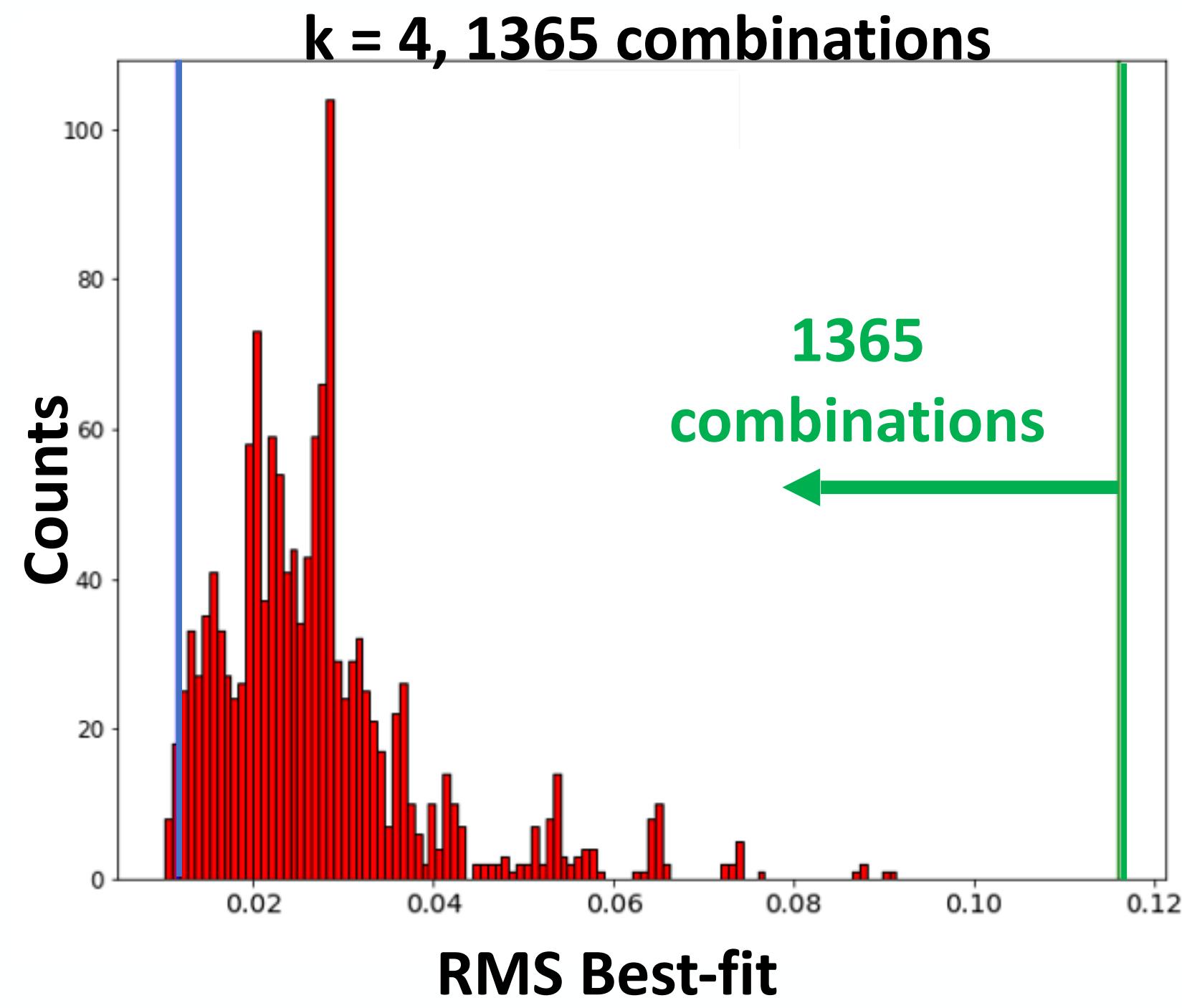
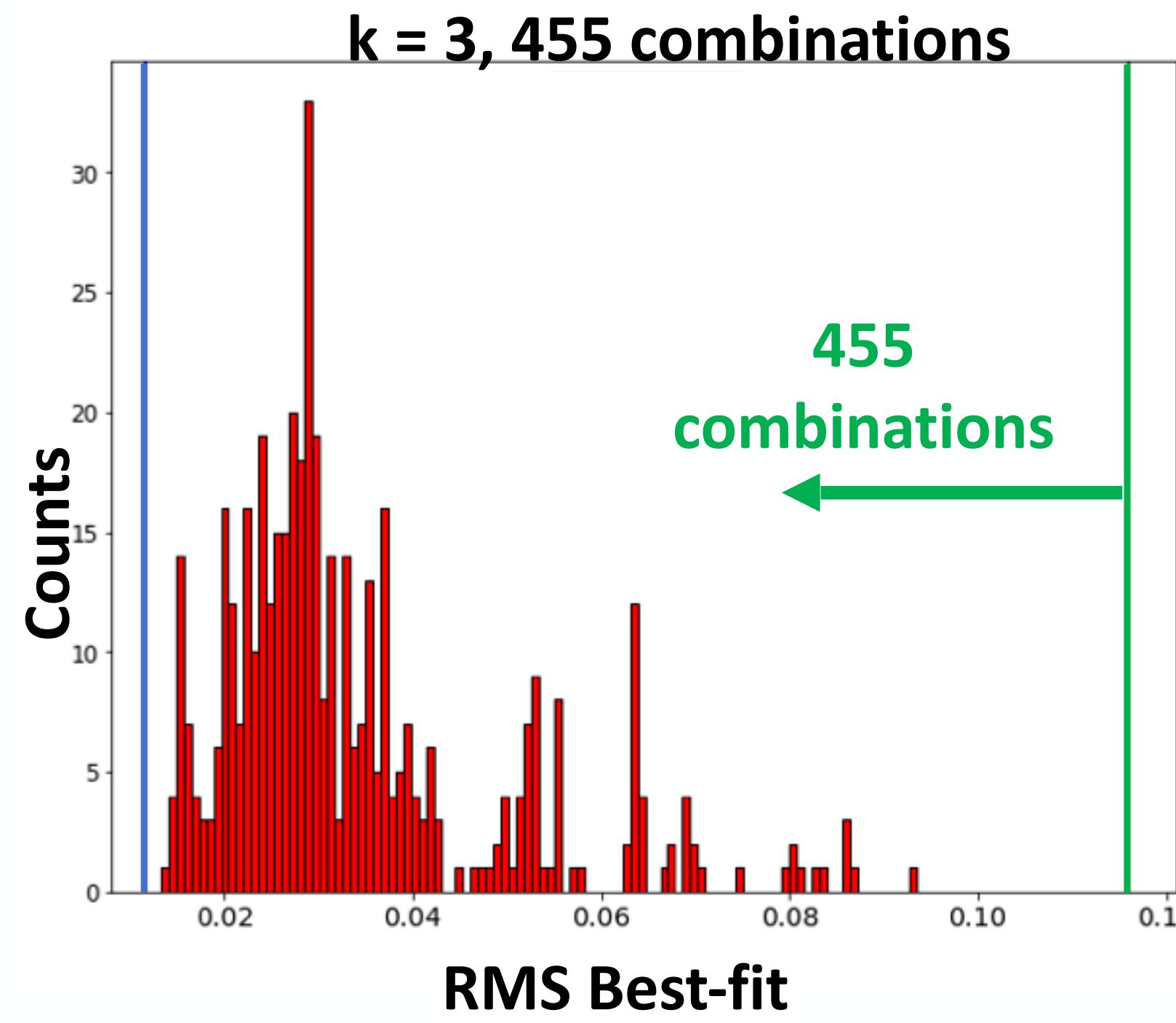
All best-fit

- 2 scenarios:
 - SNR of 5 (20% uncertainties)
 - SNR of 50 (2% uncertainties)



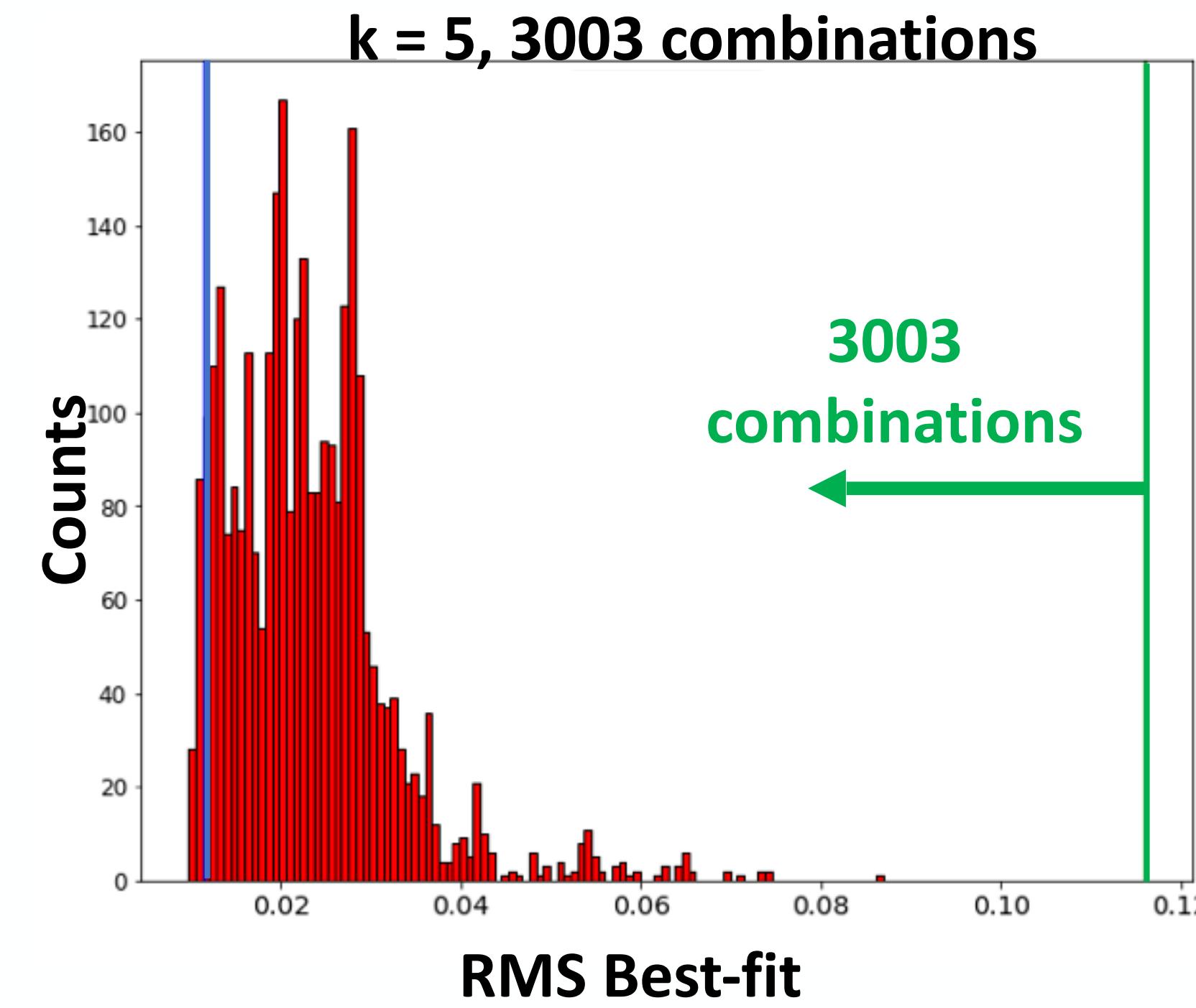
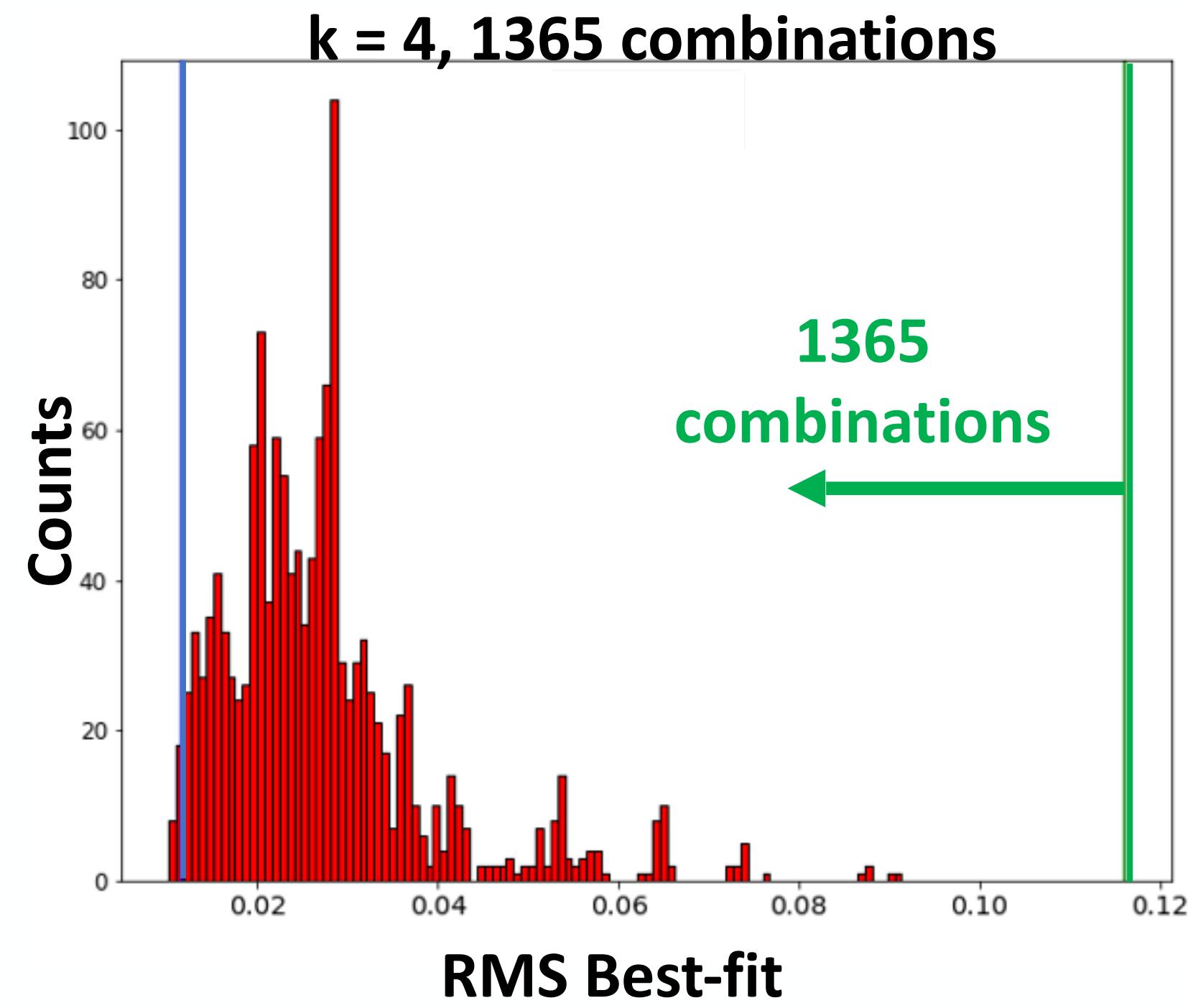
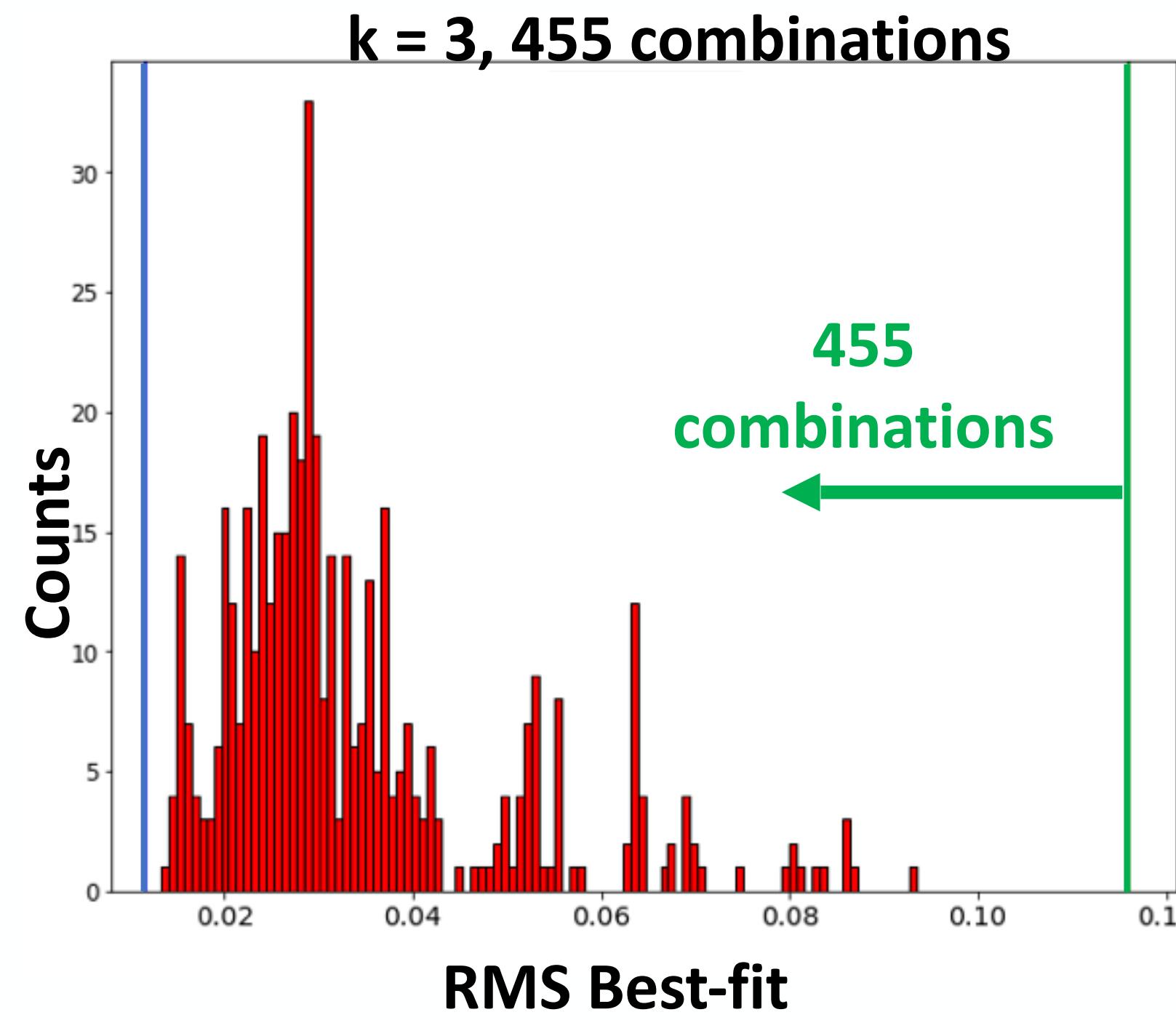
All best-fit

- 20% uncertainties: All acceptable



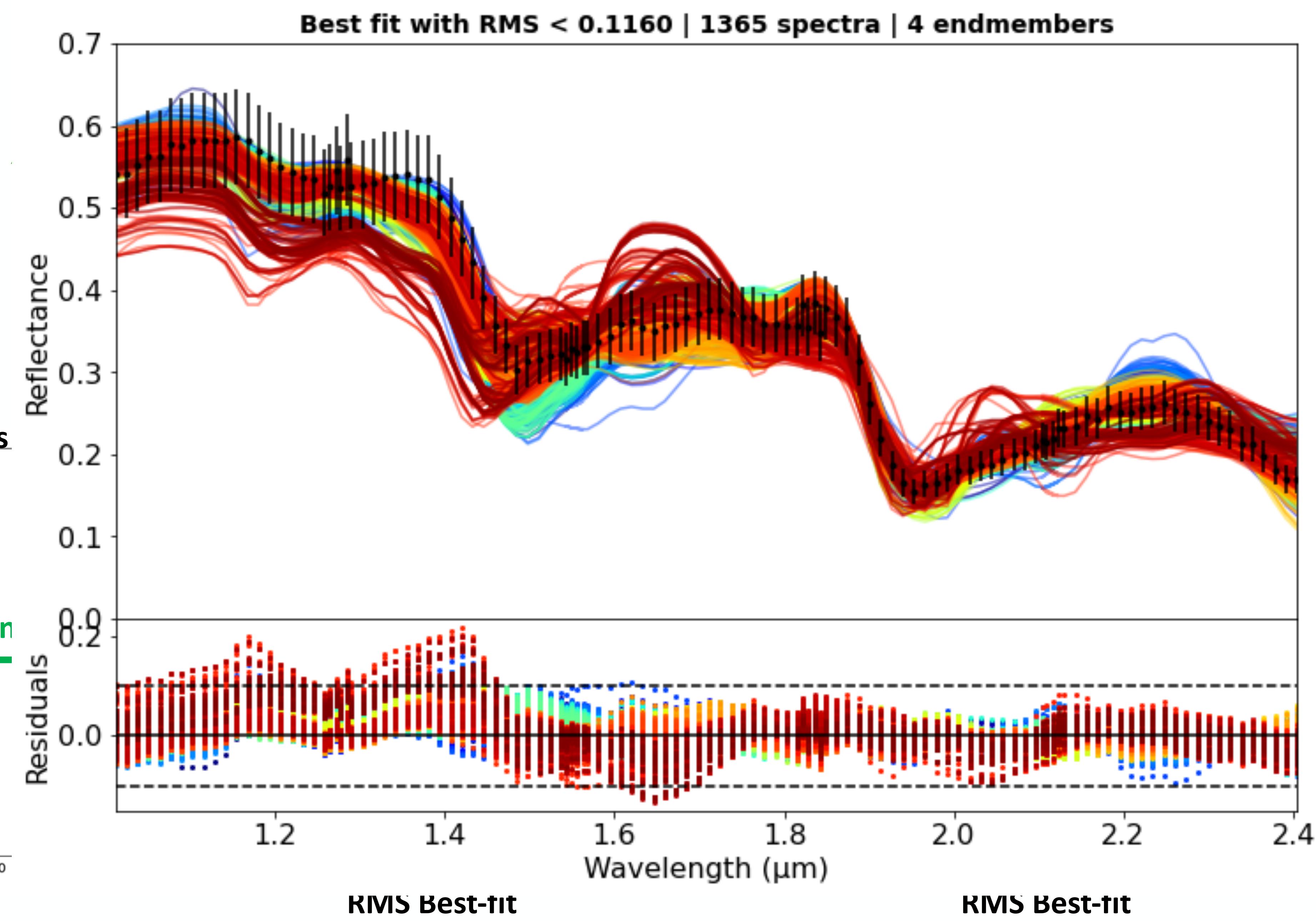
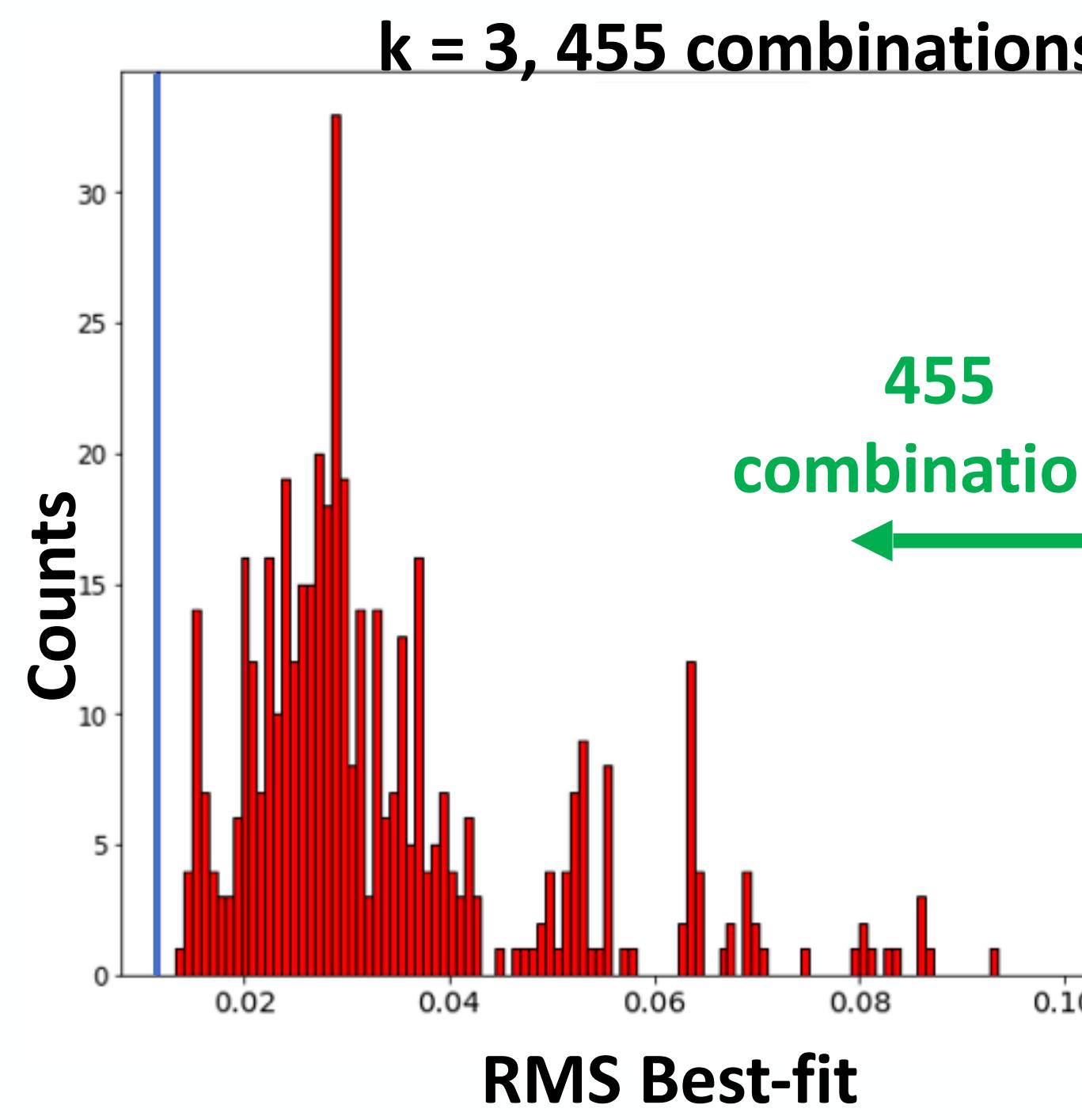
All best-fit

- 20% uncertainties: All acceptable



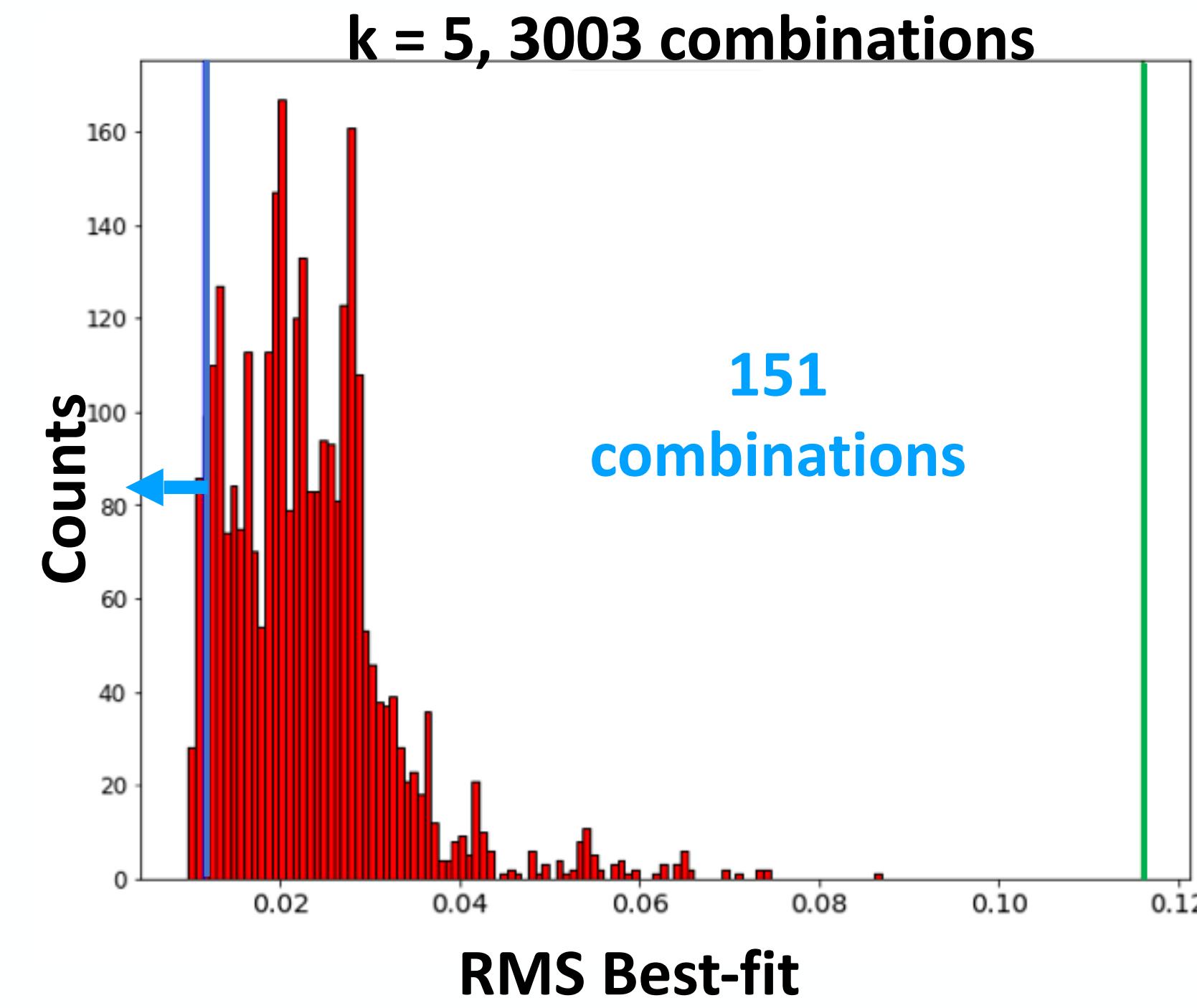
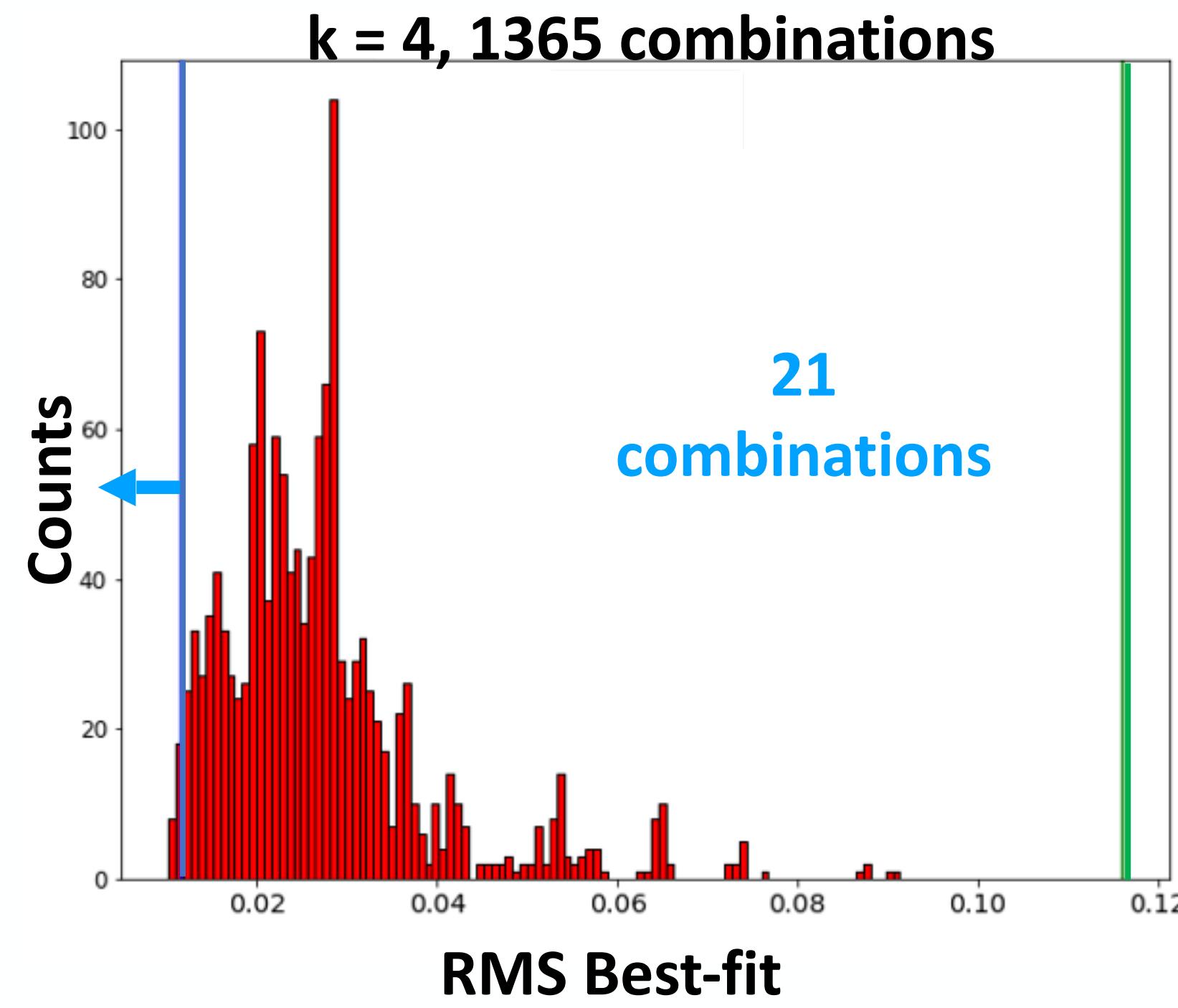
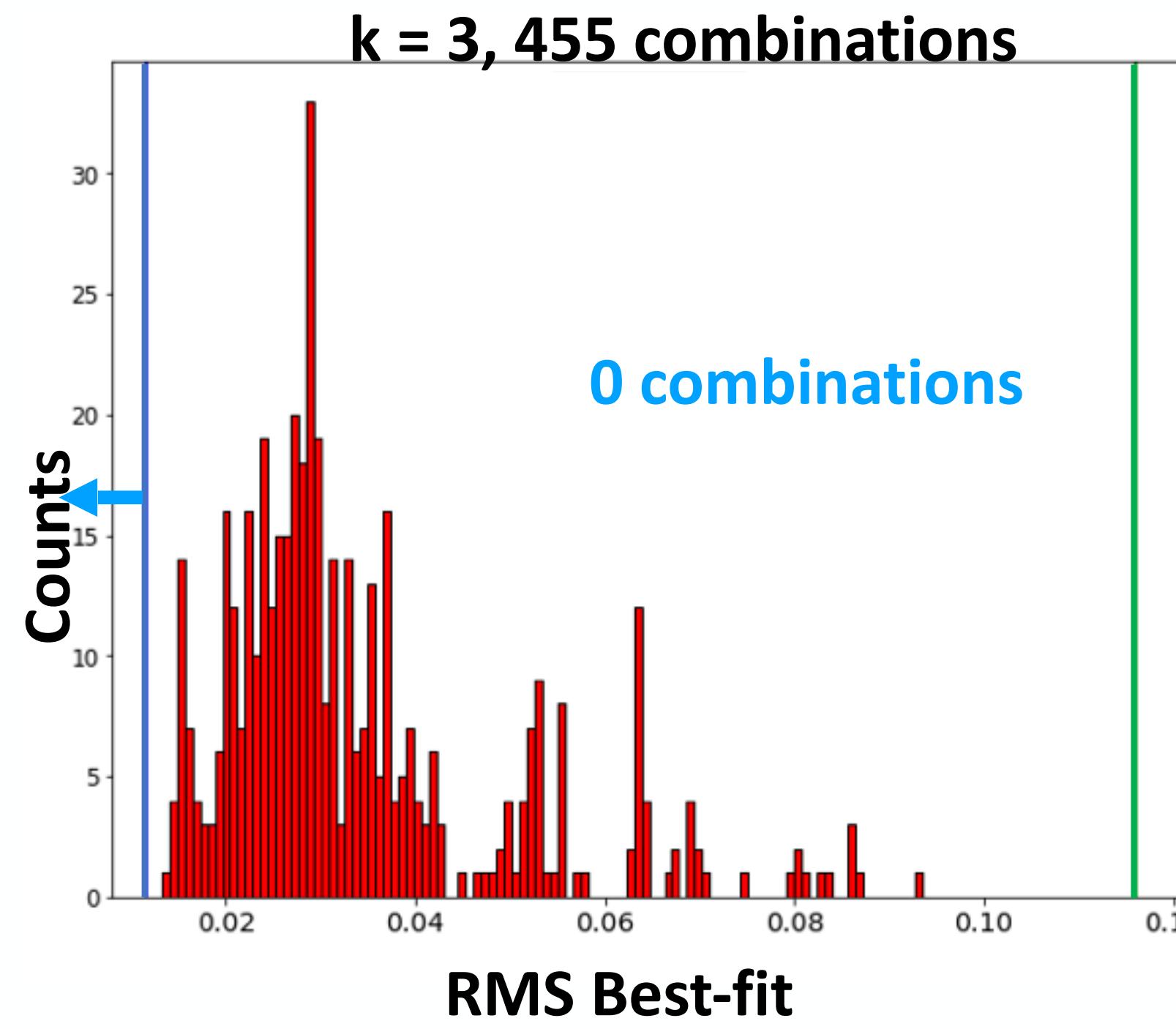
All best-fit

- 20% uncertainties:



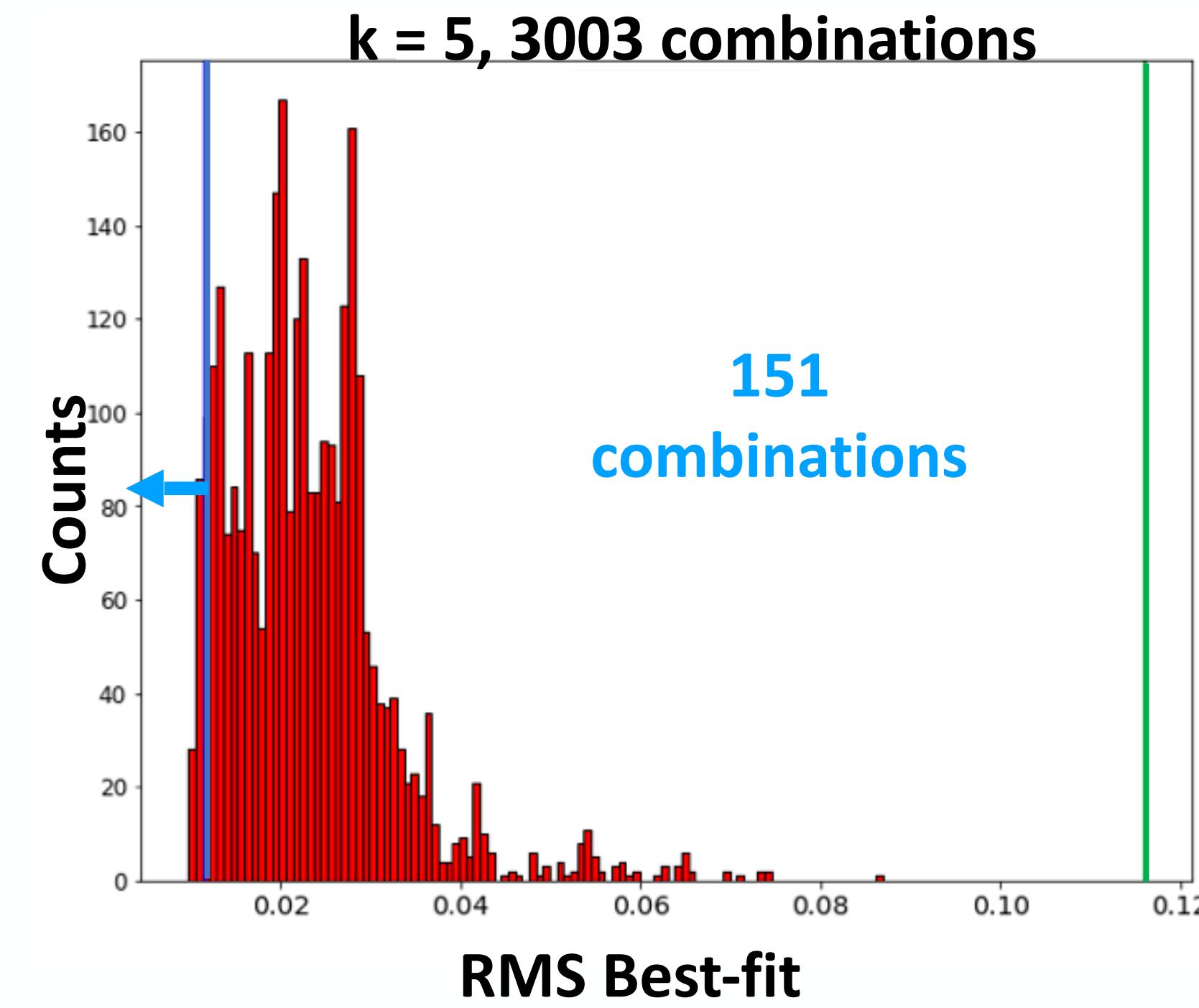
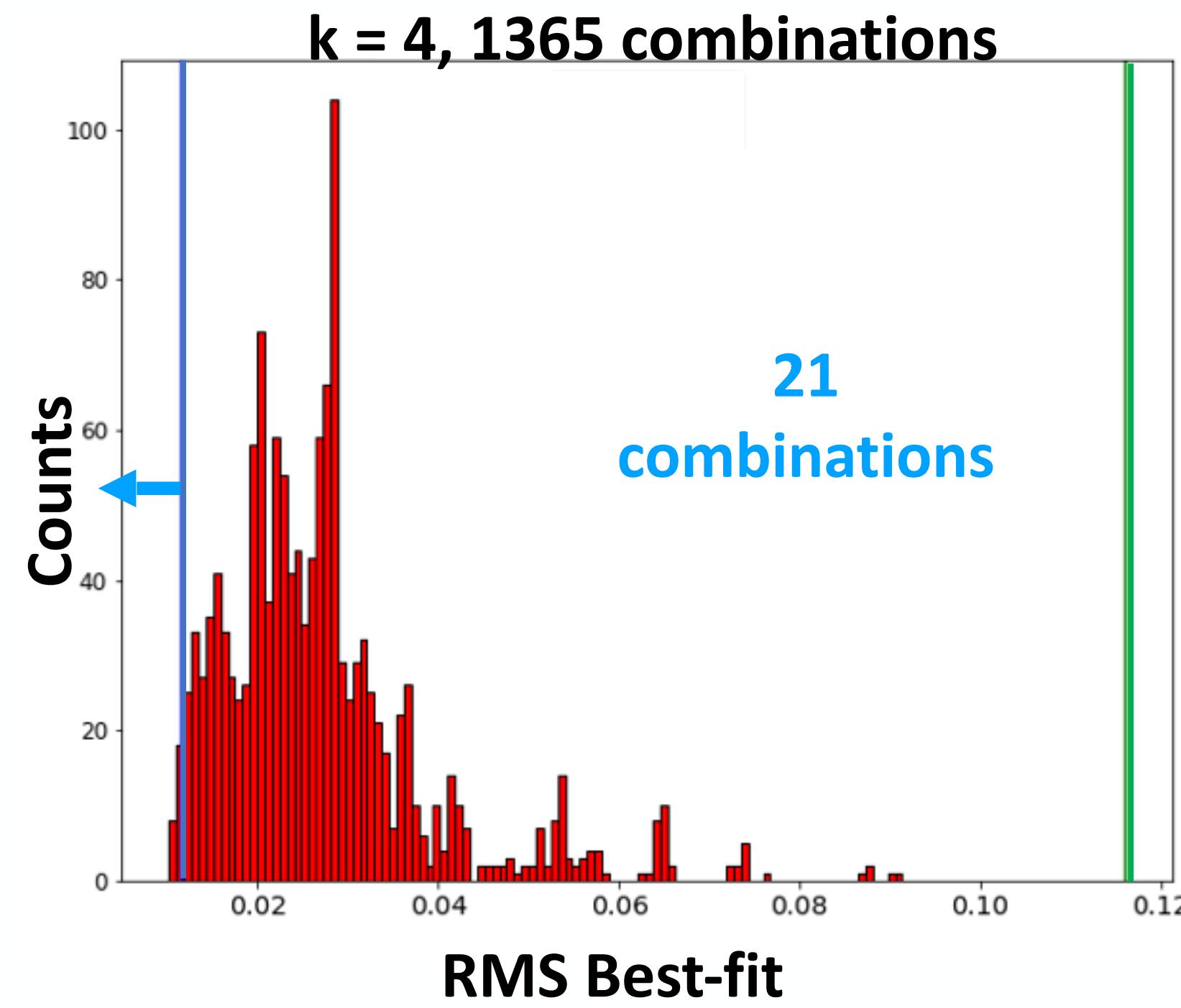
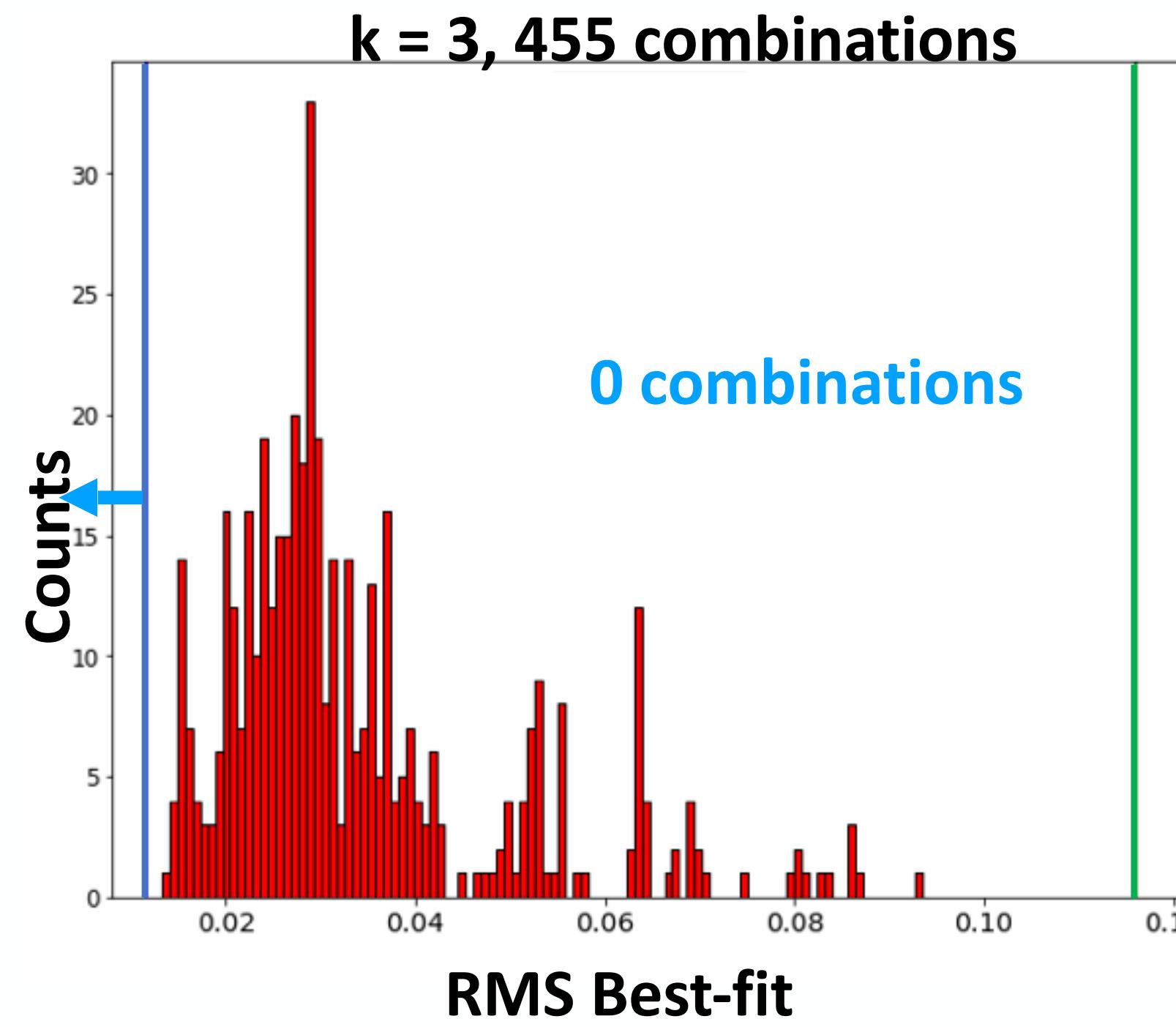
All best-fit

- 2% uncertainties: **few acceptable !**



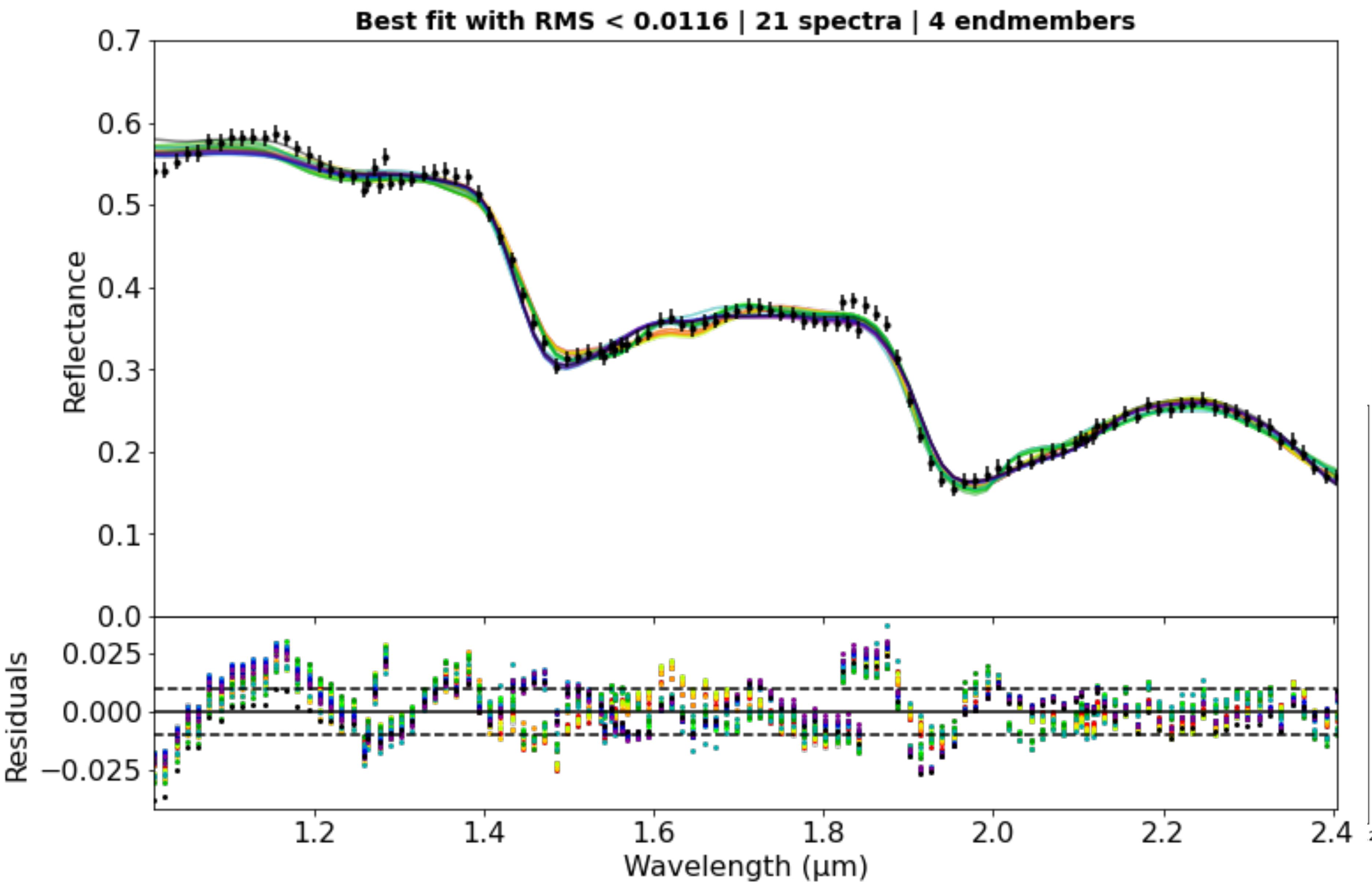
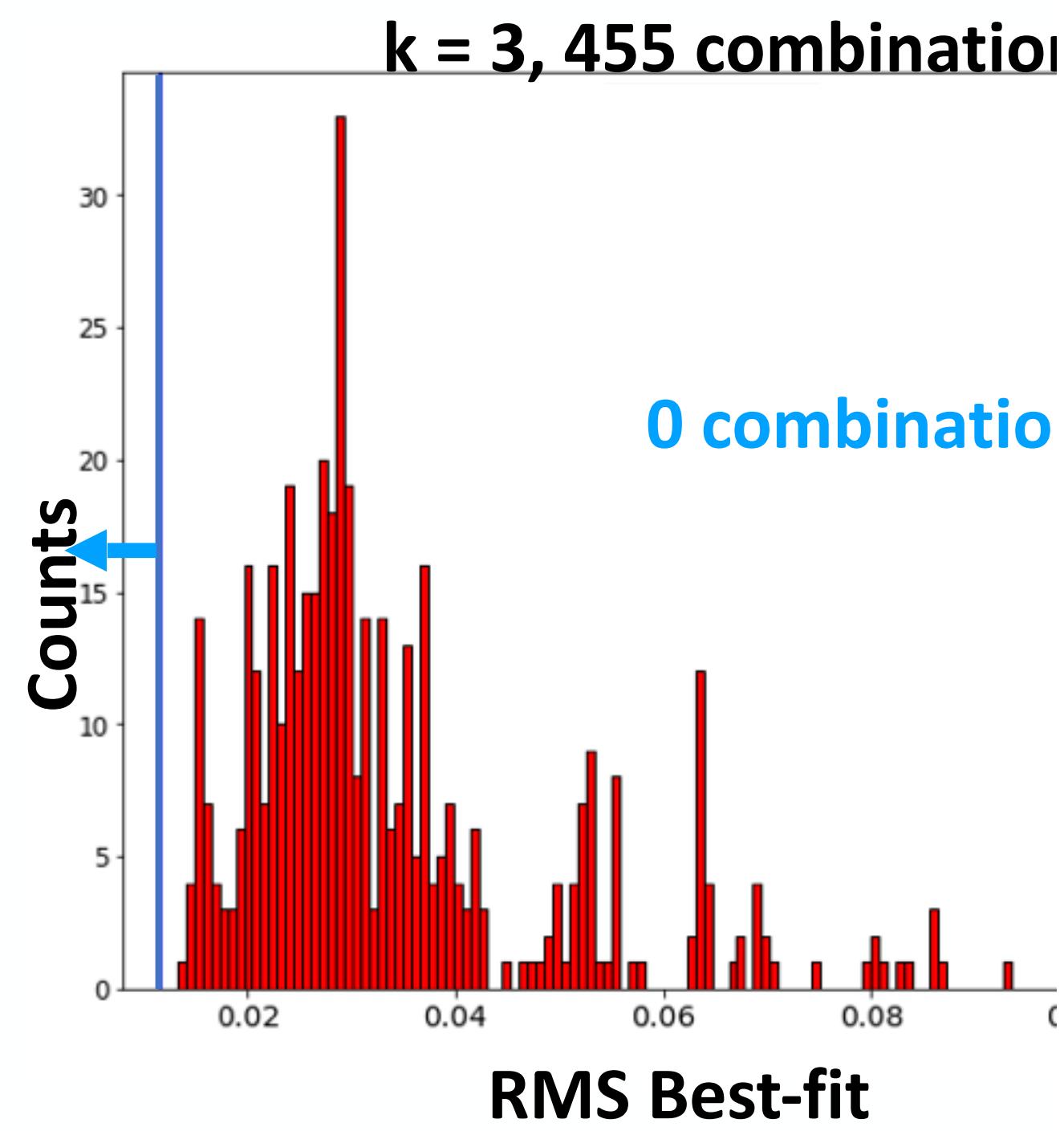
All best-fit

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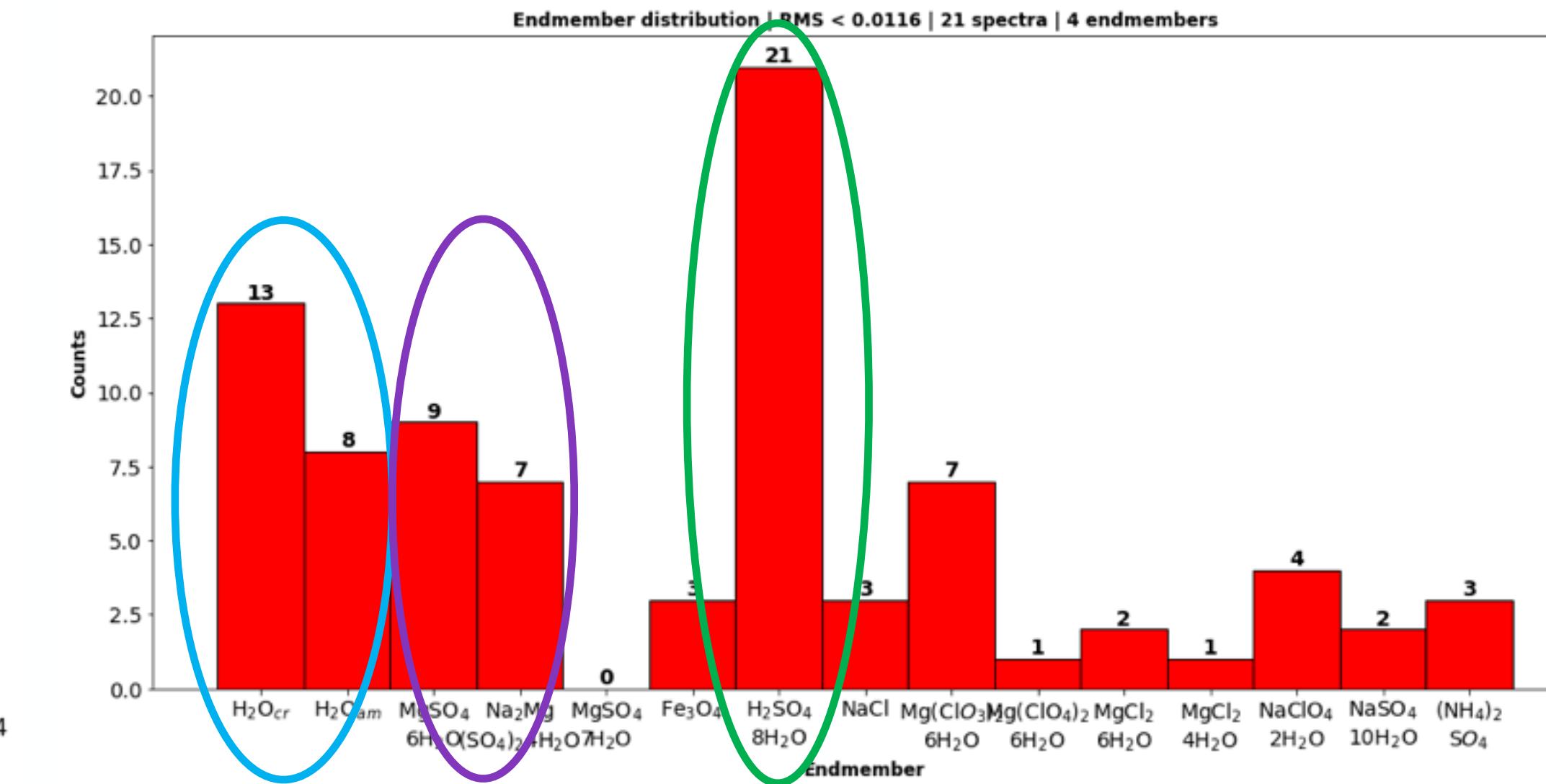
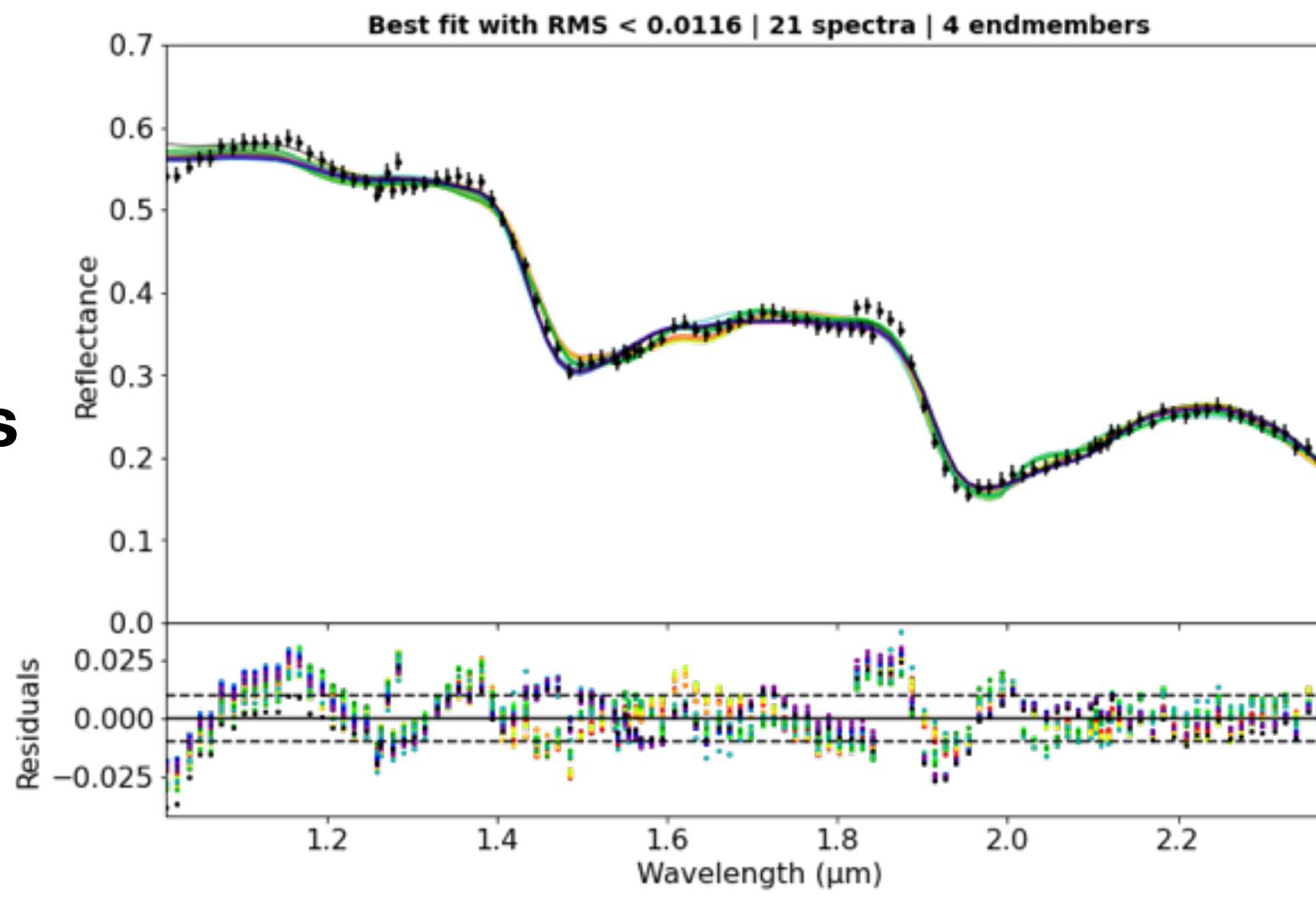
All best-fit

- 2% uncertainties:



2% uncertainties scenario

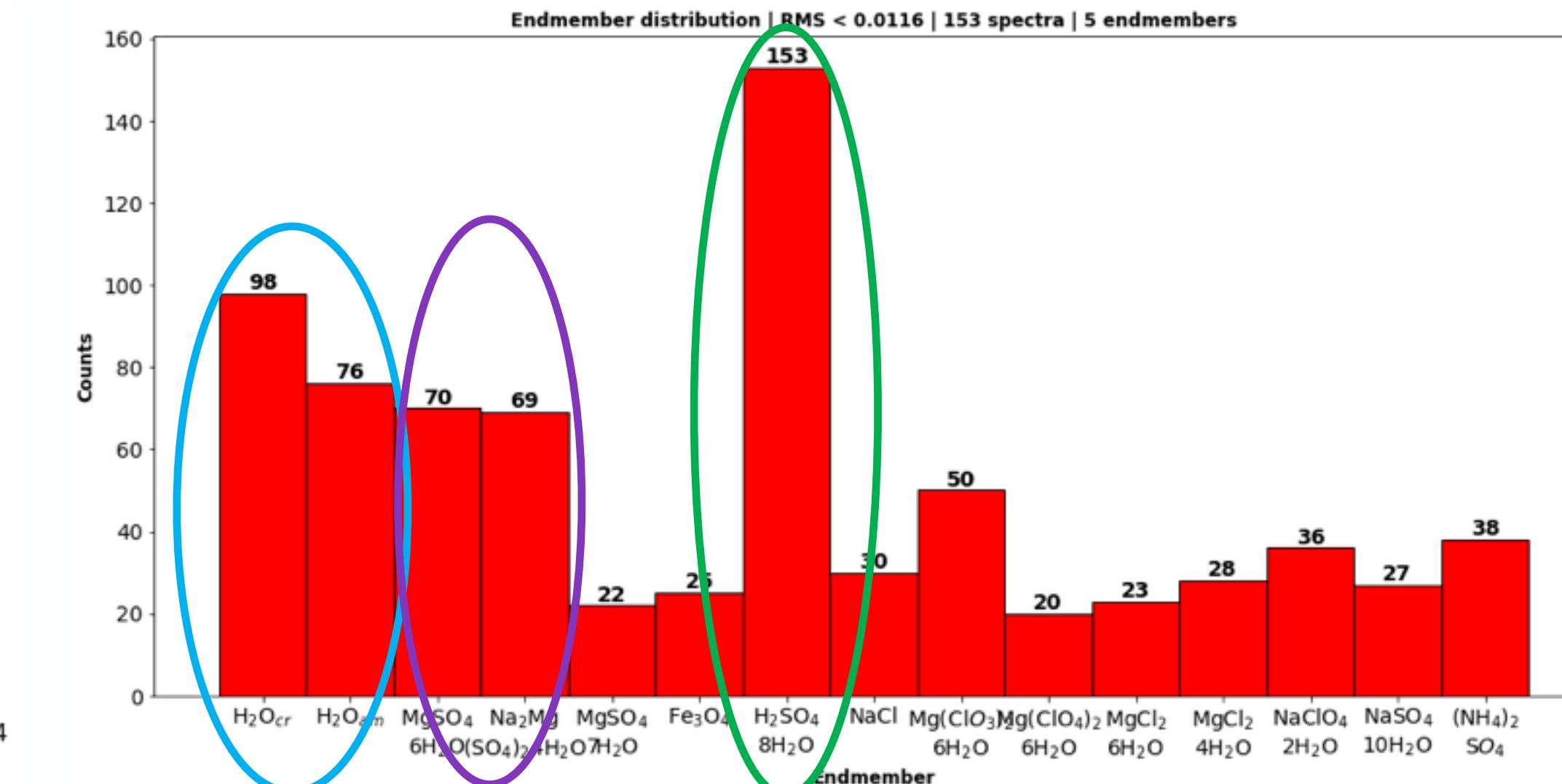
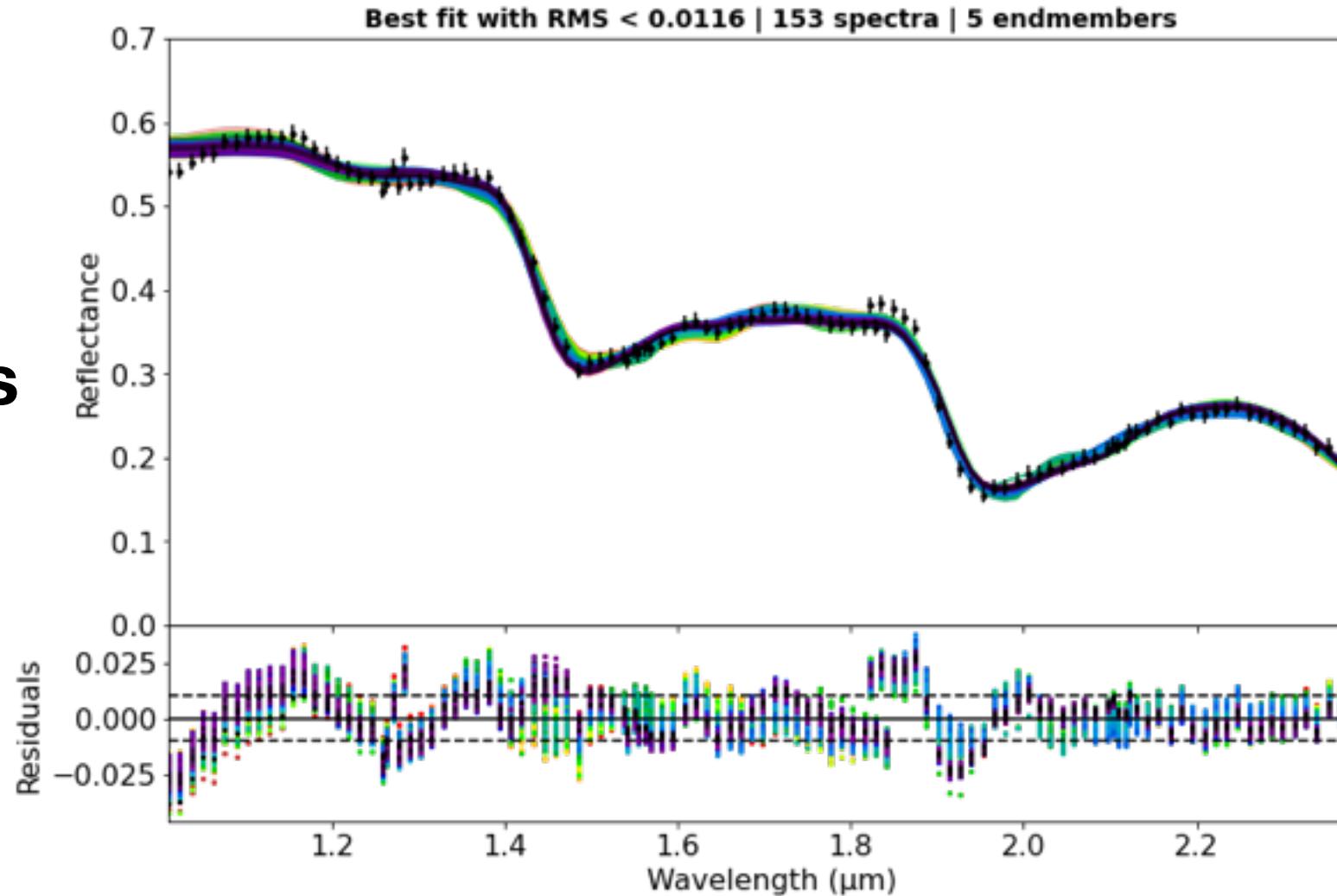
4 Compounds



Sulfuric Acid Octahydrate

Water Ice

5 Compounds



Hexahydrite Bloedite

Selection criteria

- RMS comparison → Goodness of fit per end member
- Spectral improvement factor → Spectral contribution per endmember
- Endmembers distribution → Statistics on all acceptable fits
- Numerical abundance → Representation of the medium

Criteria	H ₂ SO ₄	H ₂ O	MgSO ₄	Na ₂ Mg(SO ₄) ₂	H ₂ O	MgSO ₄	(NH ₄) ₂ SO ₄	Na ₂ SO ₄	Mg(ClO ₃) ₂	Mg(ClO ₄) ₂	MgCl ₂	NaCl	MgCl ₂	NaClO ₄	Fe ₃ O ₄
	8H ₂ O	(cr)	.6H ₂ O	.4H ₂ O	(am)	7H ₂ O		10H ₂ O	6H ₂ O	6H ₂ O	6H ₂ O	0	4H ₂ O	2H ₂ O	
RMS (SNR 5)	1	1	0.5	0.5	0.5	0.5	0.5	0	0	0	0	0	0	0	0
Improvement factor (SNR 5)	1	1	1	1	1	1	1	0.5	0	0	0	0	0	0	0
Distribution (SNR 50)	1	0.5	0.5	0.5	0.5	0	0	0	0.5	0	0	0	0	0	0
Numerical Abundances (SNR 50)	1	0	0.5	0.5	0	0.5	0	0.5	0.5	0.5	0.5	0.5	0	0	0
Average	1	0.625	0.625	0.625	0.5	0.5	0.375	0.25	0.25	0.125	0.125	0.125	0	0	0

Conclusion: surface composition

- Very good spectral fit to NIMS data without artificial compounds and/or high porosity
- Water ice & Sulfuric Acid Octahydrate:
Mandatory
- Hydrated Sulfates & Chlorinates:
Indistinguishable but **required**
- Magnetite, Magnesium Chloride, Sodium perchlorate: **not necessary** or **absent**

Conclusion: surface composition

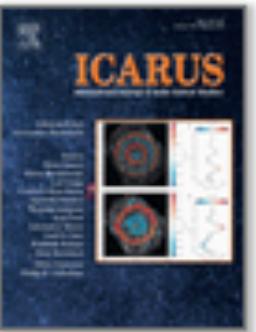
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In Press, Journal Pre-proof [?](#)



Research Paper

Selection of chemical species for Europa's surface using Galileo/NIMS

G. Cruz Mermyn^a , F. Schmidt^{a, b}, T. Cornet^c, I. Belgacem^d, N. Altobelli^d

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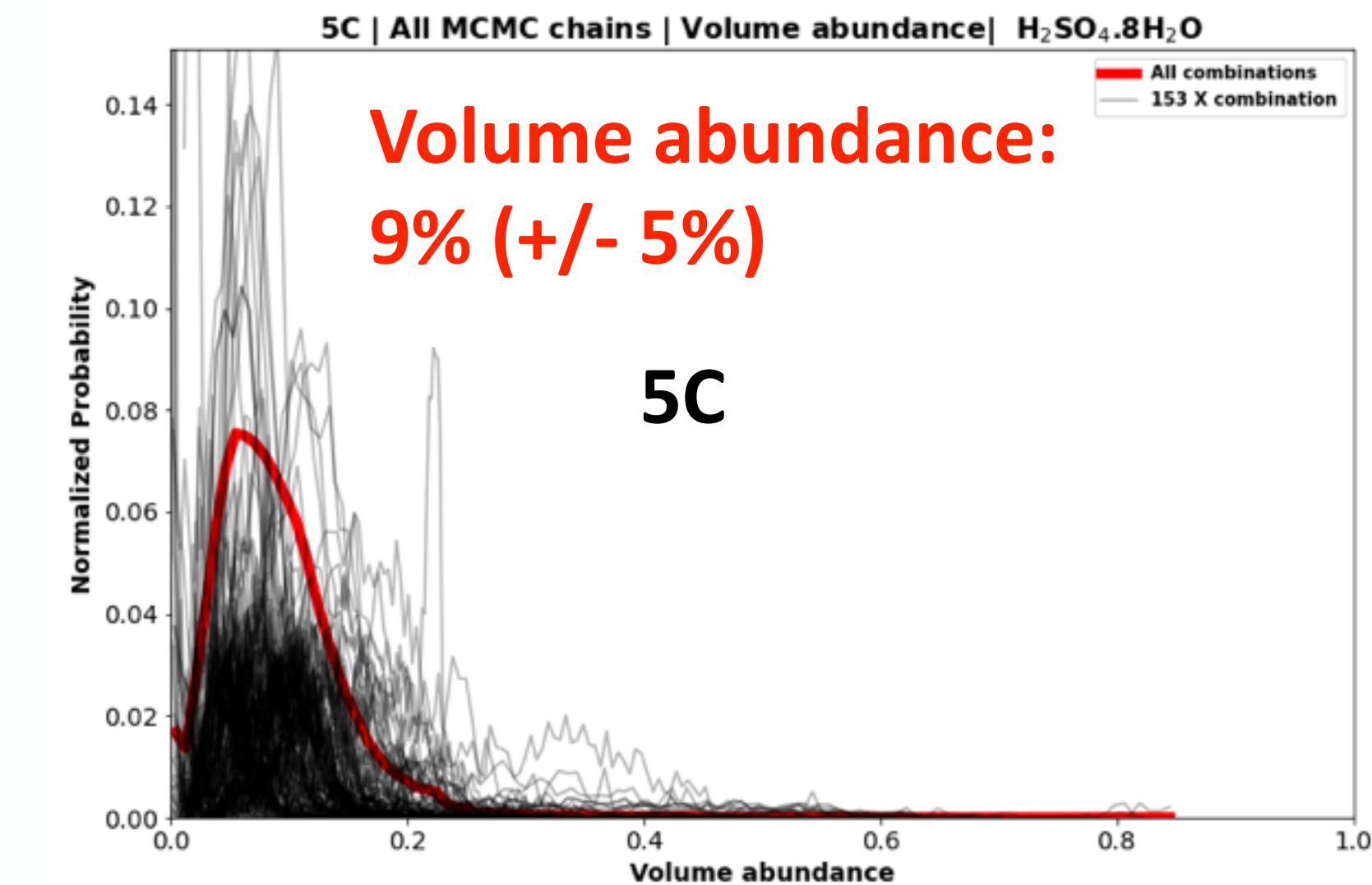
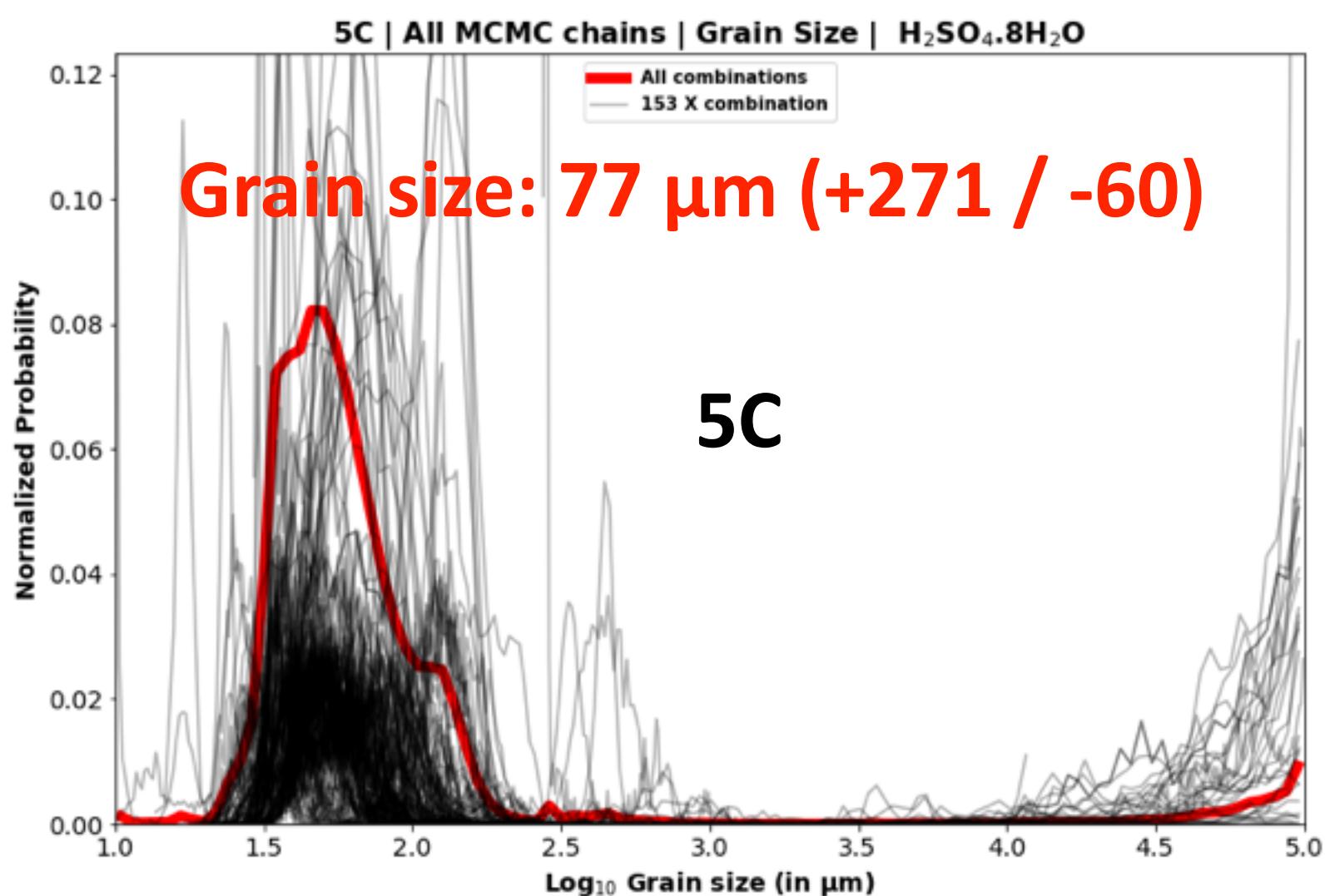
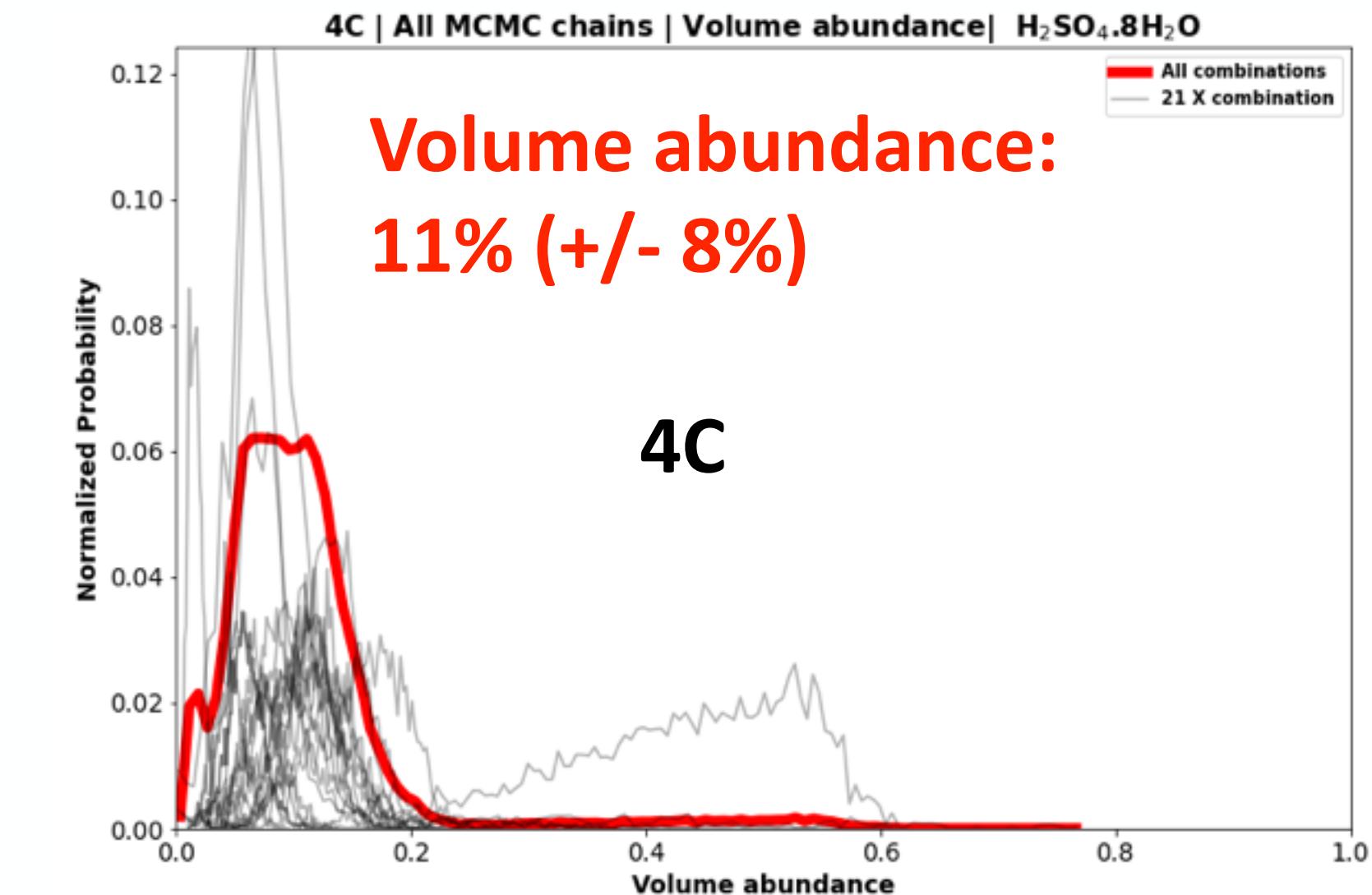
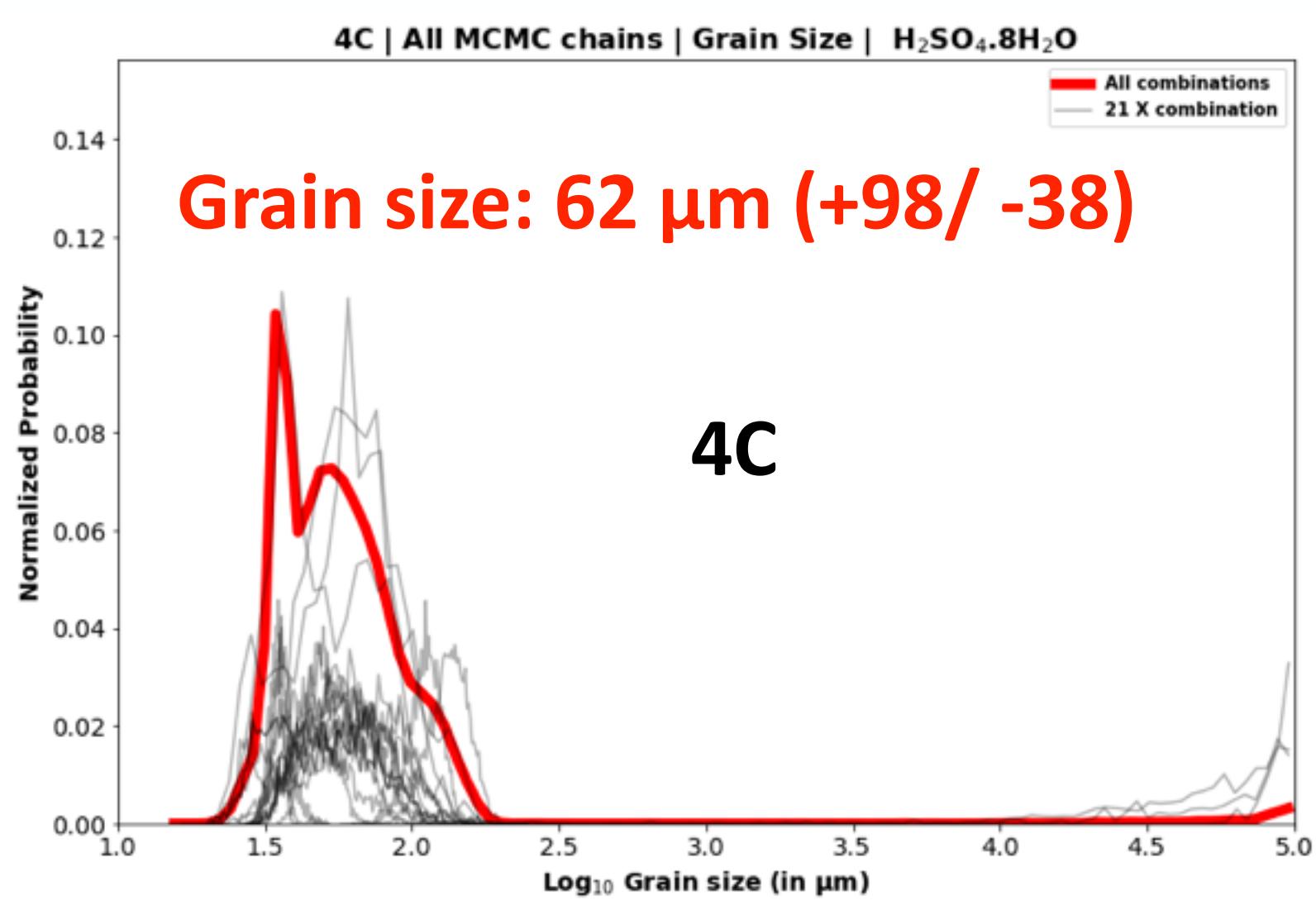
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Microphysics: grain size & abundance

- All 21 and 153 acceptable combinations
- MCMC modeling with 2% uncertainties
- Probability Density Function



Microphysics: surface roughness

